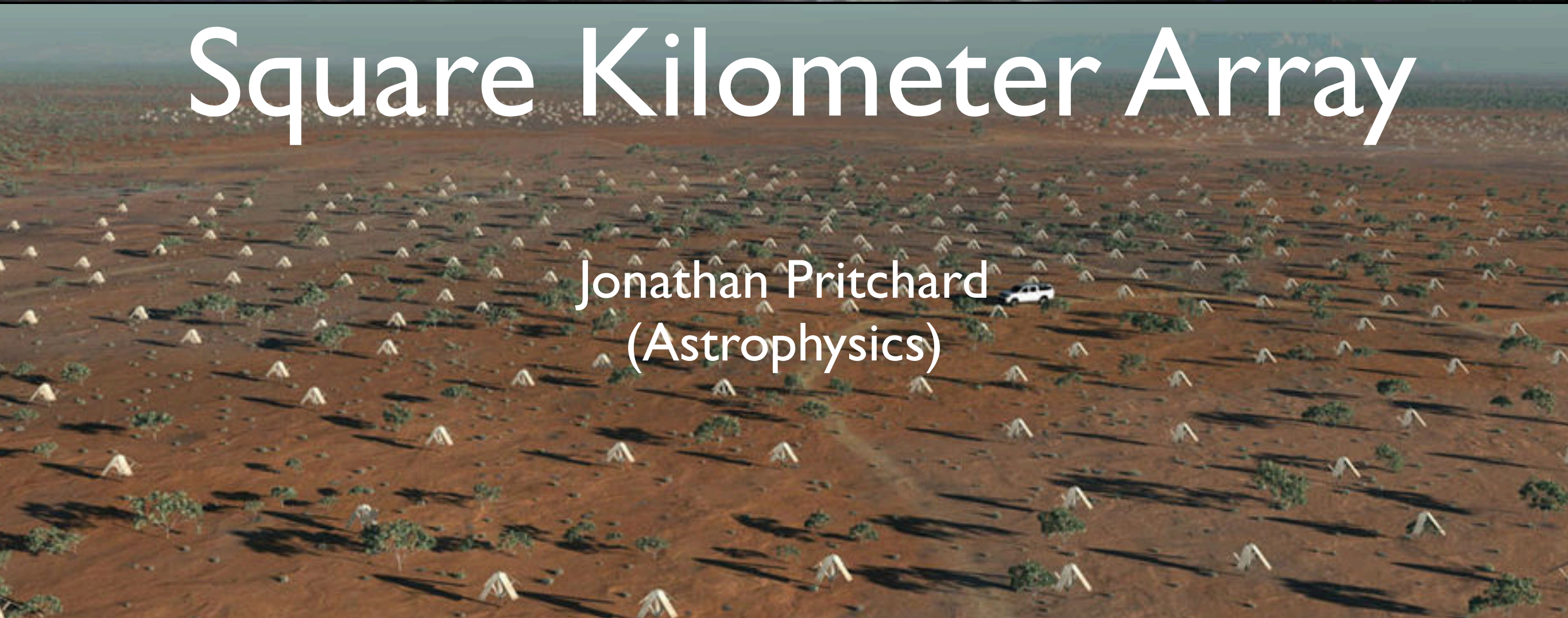


In the beginning of the Dark Ages, electrically neutral hydrogen gas filled the universe. As stars formed, they ionized the regions immediately around them, creating bubbles here and there. Eventually these bubbles merged together, and intergalactic gas became entirely ionized.

# Cosmic Dawn &

# Square Kilometer Array



Jonathan Pritchard  
(Astrophysics)



**RADIO QUIET ZONE**



**SWITCH OFF  
MOBILE PHONES**

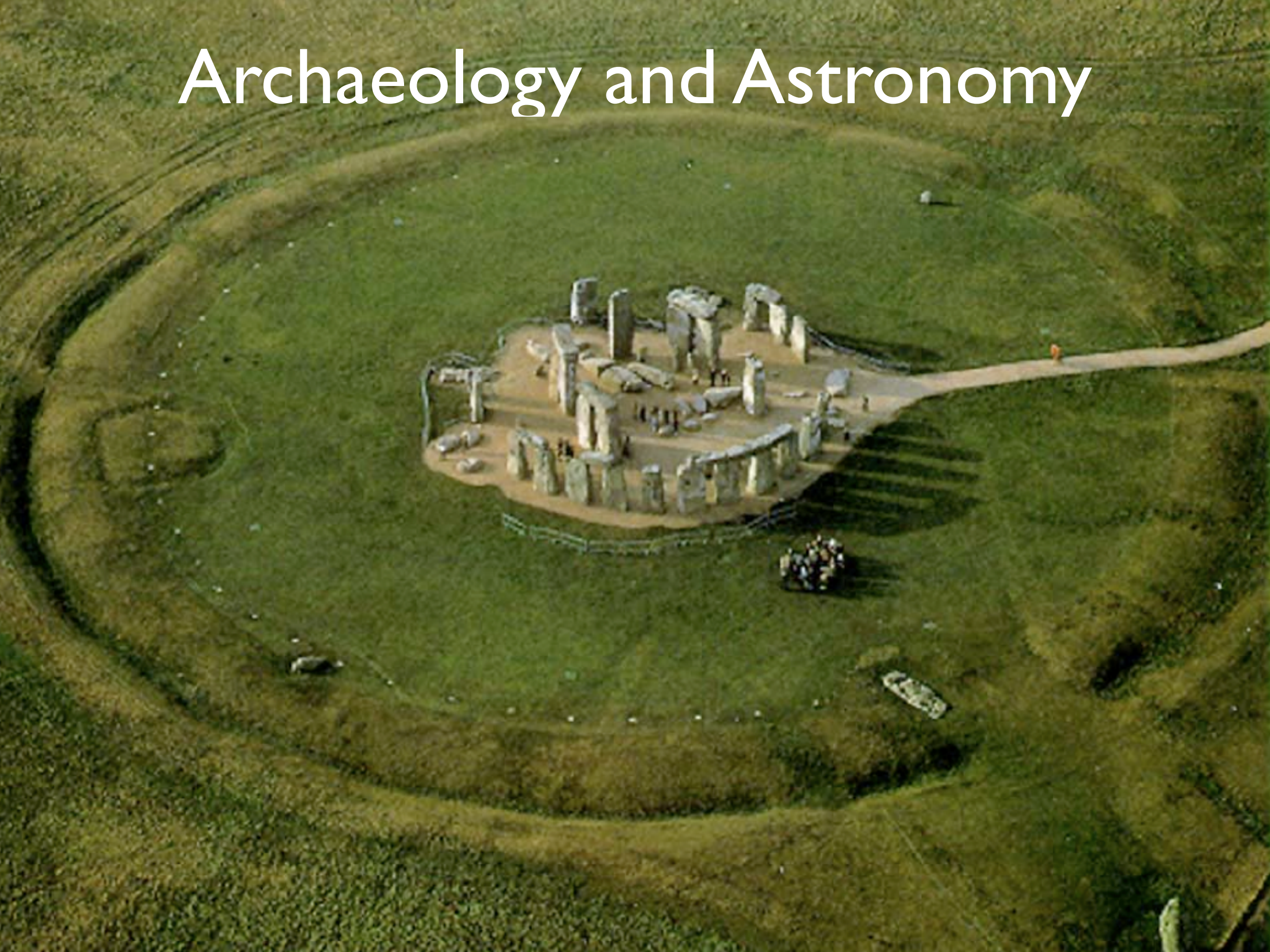








# Archaeology and Astronomy





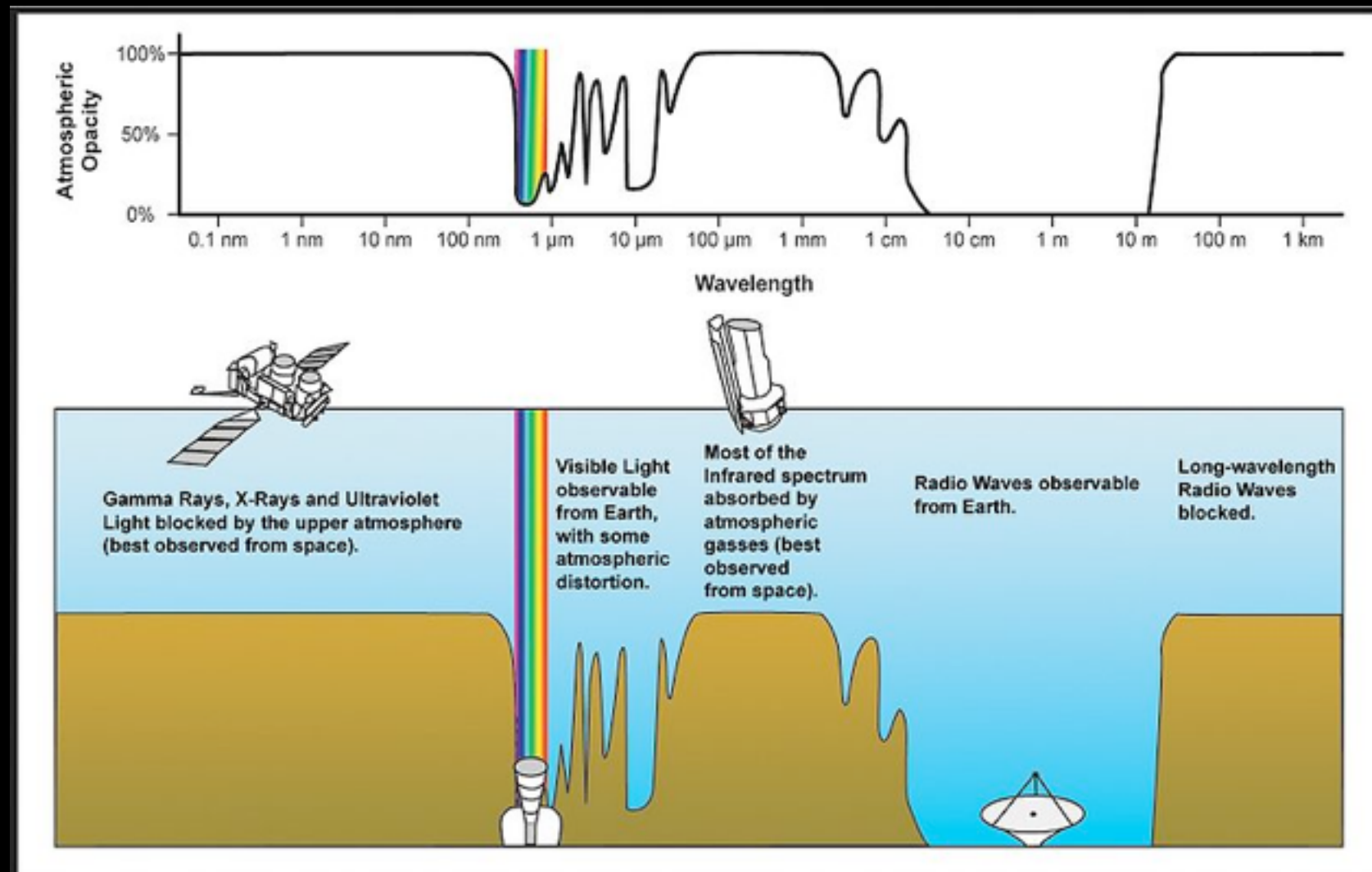
# Seeing the past

- Light travels at 300,000 km/s. Fast, but finite.
- When we see distant objects, we're seeing them as they were when that light was emitted
- We're seeing the past
- Moon is ~1 light seconds away  
Sun is ~8 light minutes away
- We don't get to choose when we see. Each distance corresponds to a specific time in cosmic history



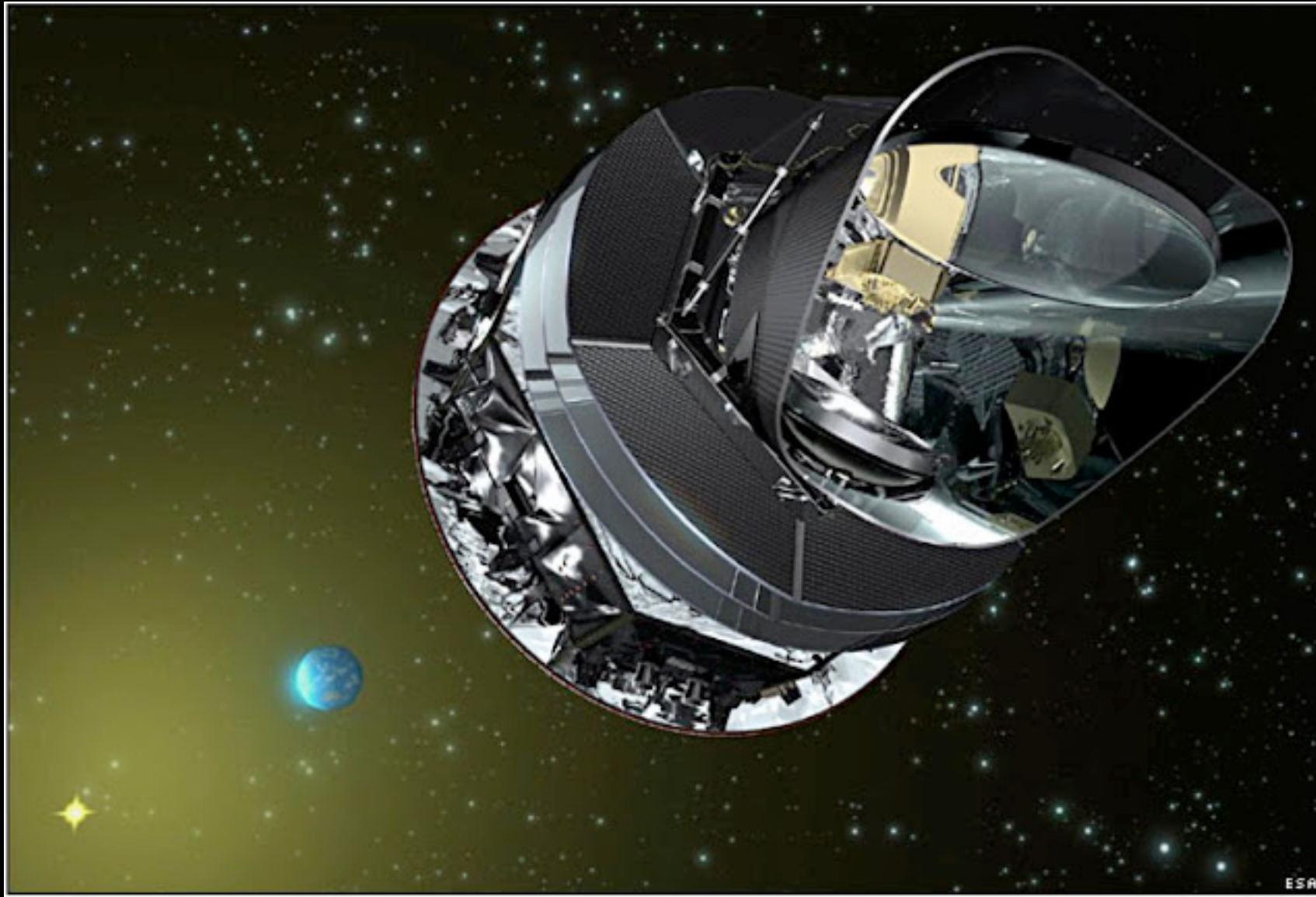
# What do we see?

- Different frequencies show us different objects at different distances.
- Limits set by technology and absorption along the line of sight





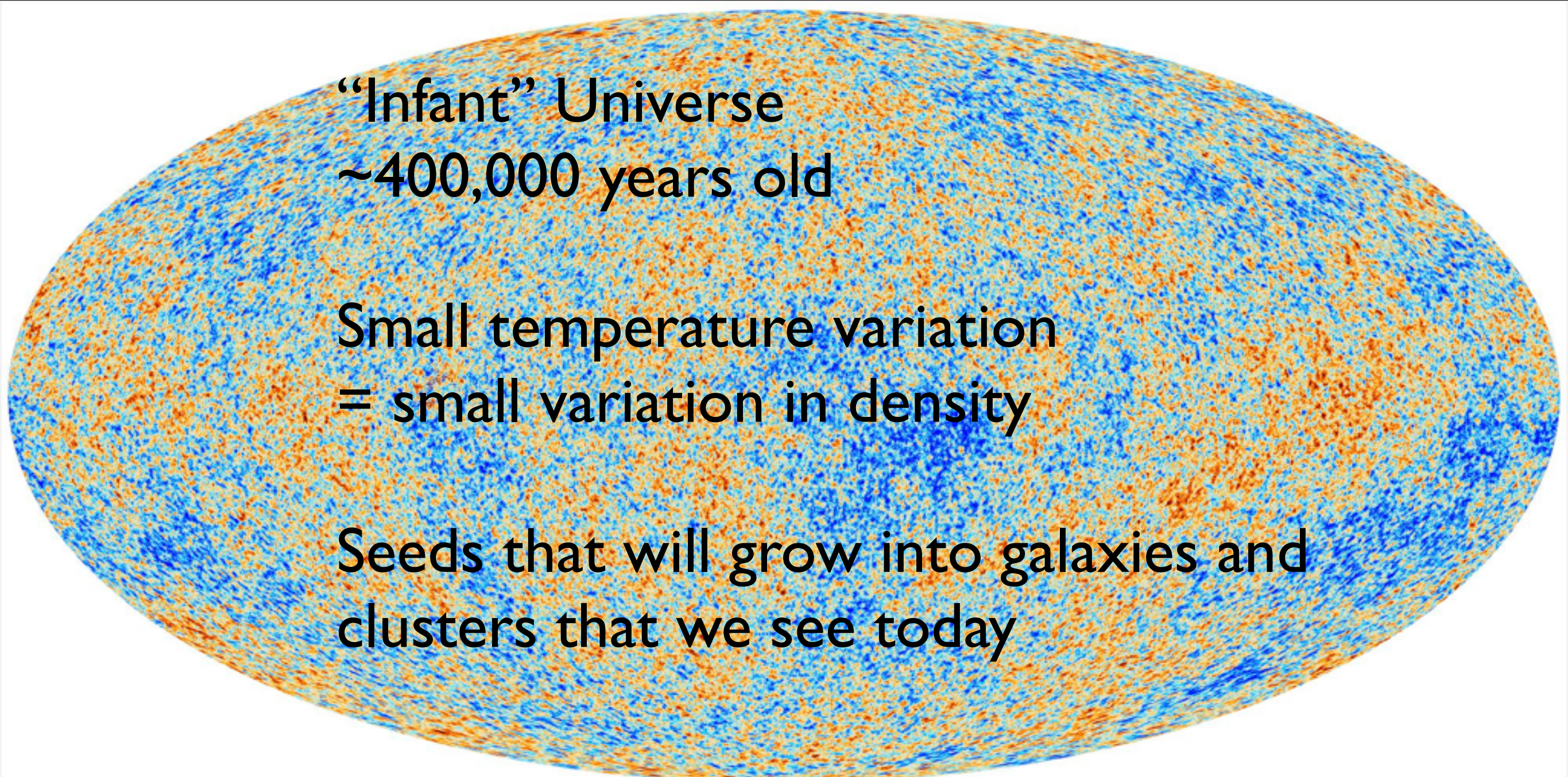
# Cosmic Microwave Background



Planck satellite



# Cosmic Microwave Background



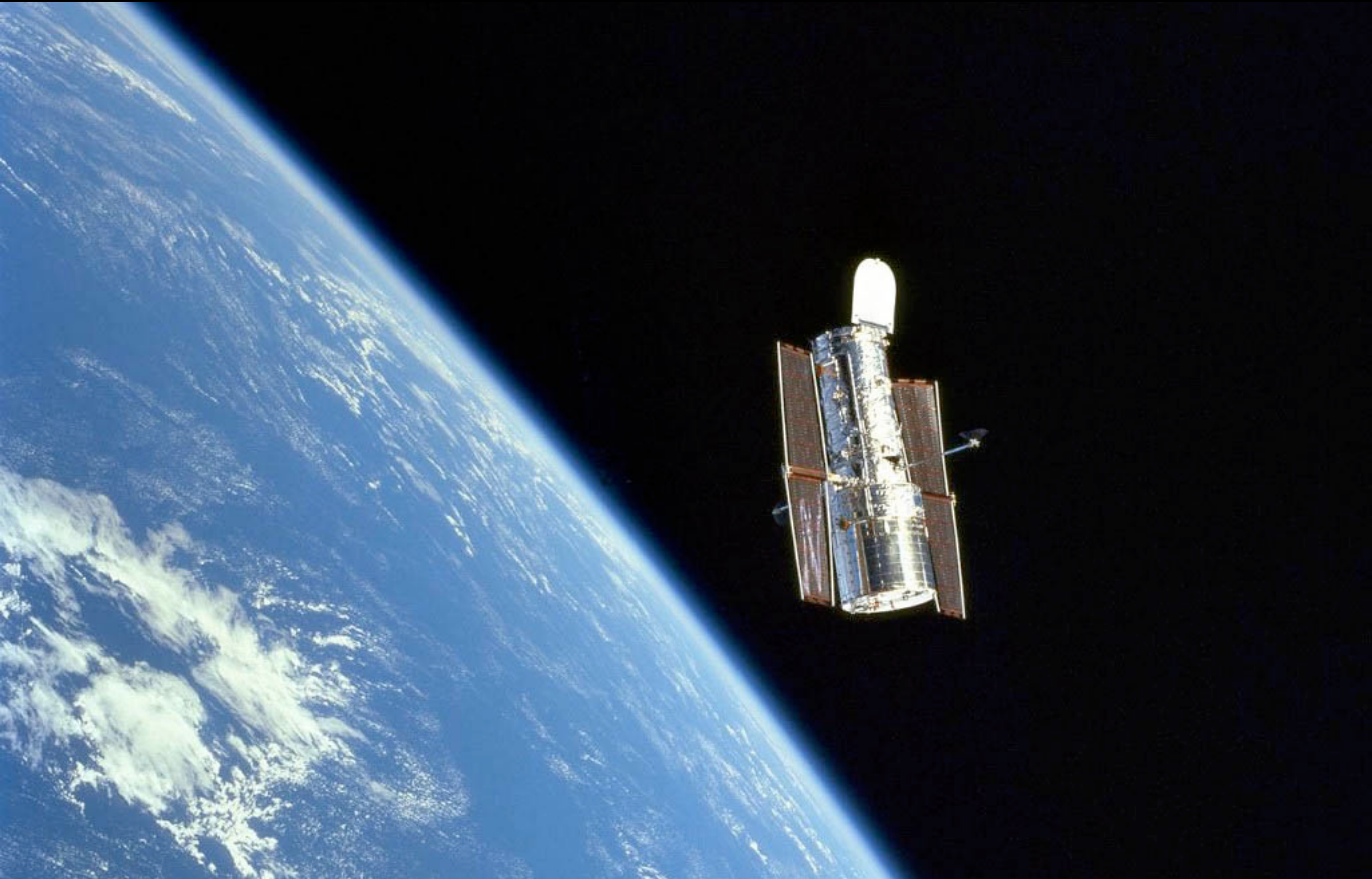
“Infant” Universe  
~400,000 years old

Small temperature variation  
= small variation in density

Seeds that will grow into galaxies and  
clusters that we see today



# Hubble Space Telescope





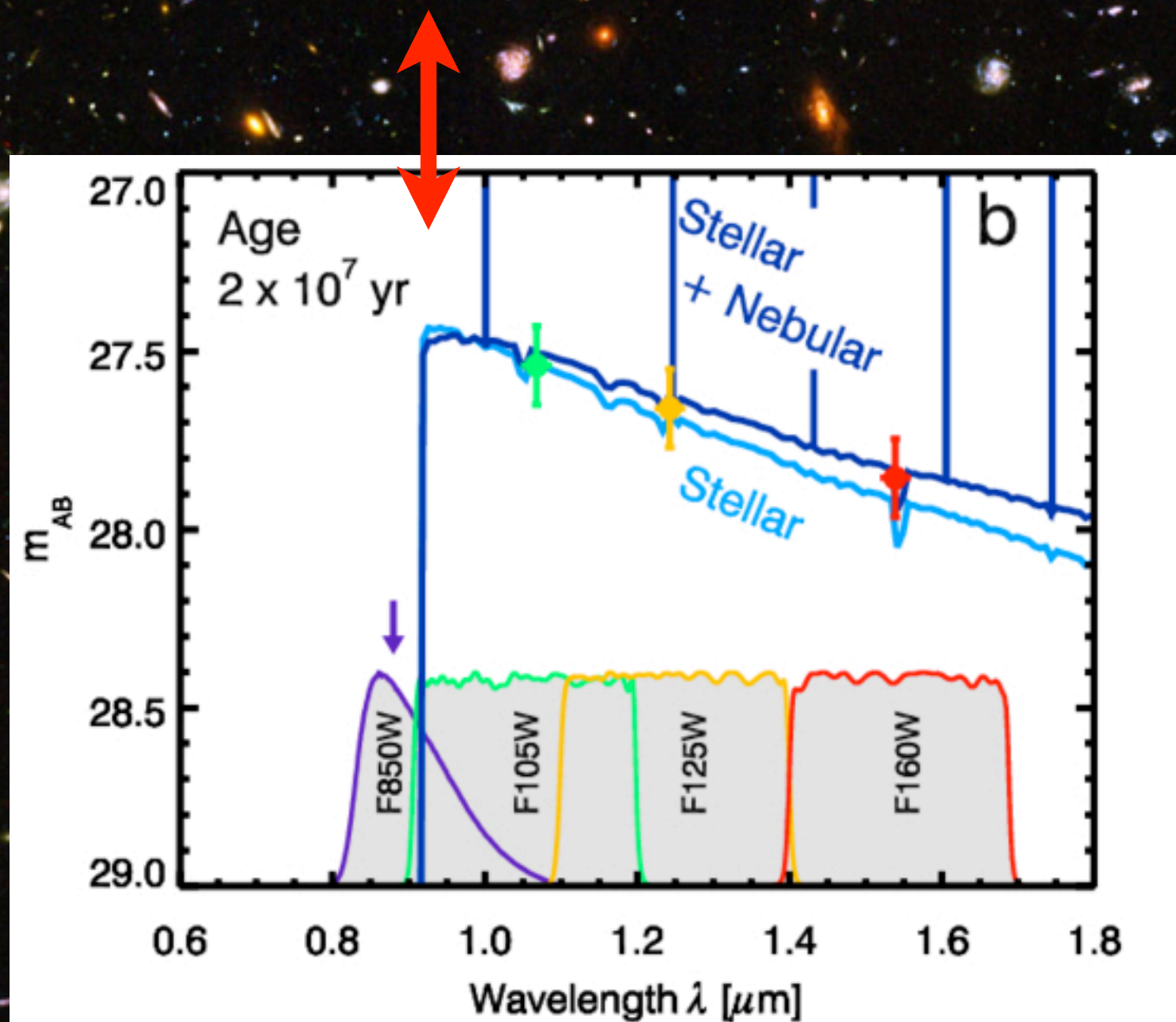
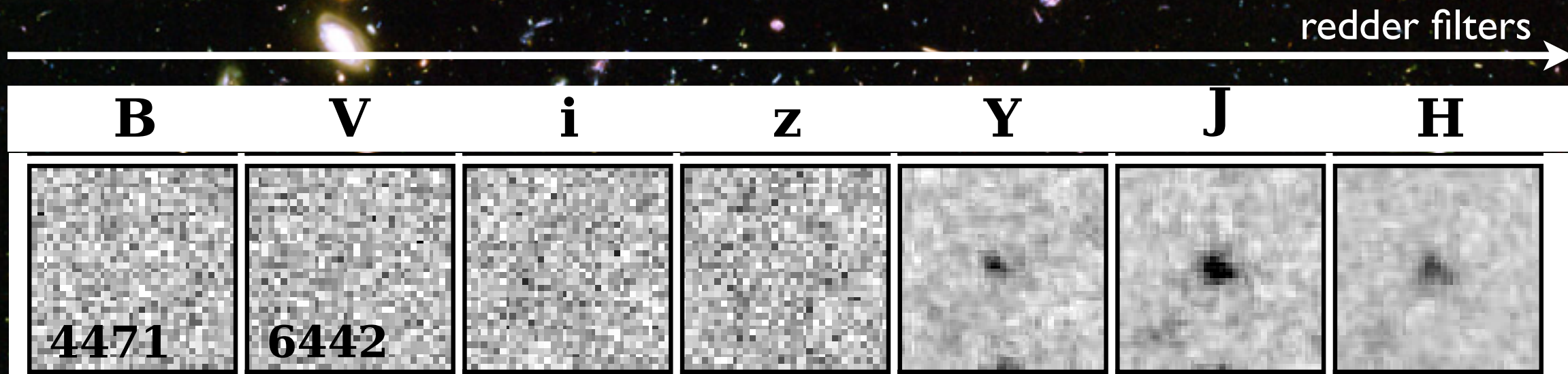
# Hubble Ultra Deep Field

The background of the slide is a vast field of galaxies, known as the Hubble Ultra Deep Field. It shows a dense collection of galaxies in various colors (blue, yellow, red, white) and shapes (spiral, elliptical, irregular), scattered across a dark cosmic background. The galaxies are of various sizes and orientations, representing a wide range of galaxy types and evolutionary stages.

- Patch smaller than the moon on sky
- Furthest galaxies are located at distances when Universe was 1 billion years old
- “Mature” Universe



# Hubble (and soon JWST) identify high redshift galaxies as “drop outs”

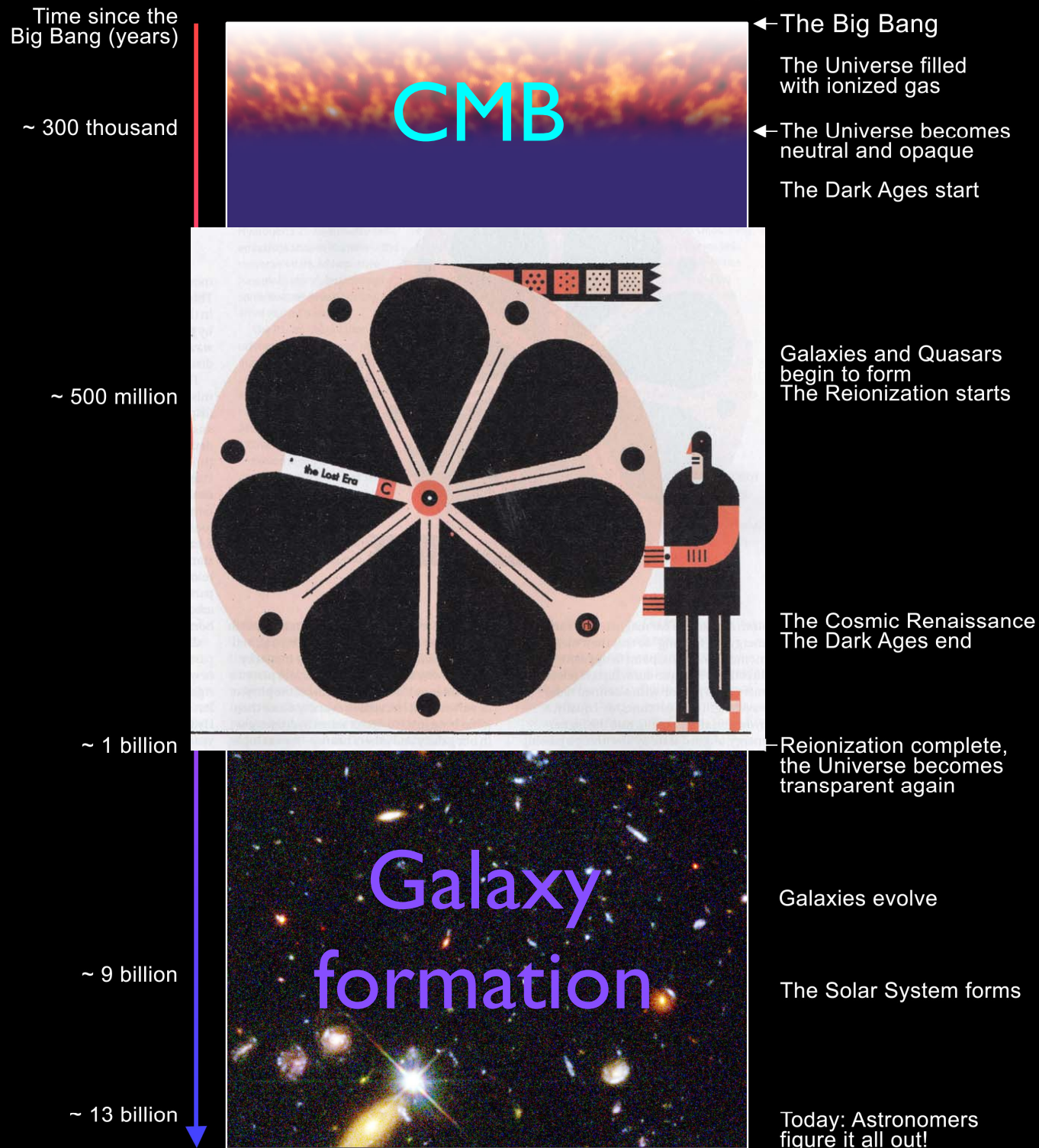




# The Cosmic Timeline

## What is the Reionization Era?

A Schematic Outline of the Cosmic History



S.G. Djorgovski et al. & Digital Media Center, Caltech

- CMB shows initial conditions
- Deep galaxy surveys show later galaxies
- What went on in between?
- Currently NO observations!





# The Missing Reel

---

**DIRECTOR** T. Universe

---

**CAMERAMAN** S.K.A.

---

<b>DATE</b>	<b>TAKE</b>	<b>SCENE</b>
13.75 GYrABB	1	$10^{51}$

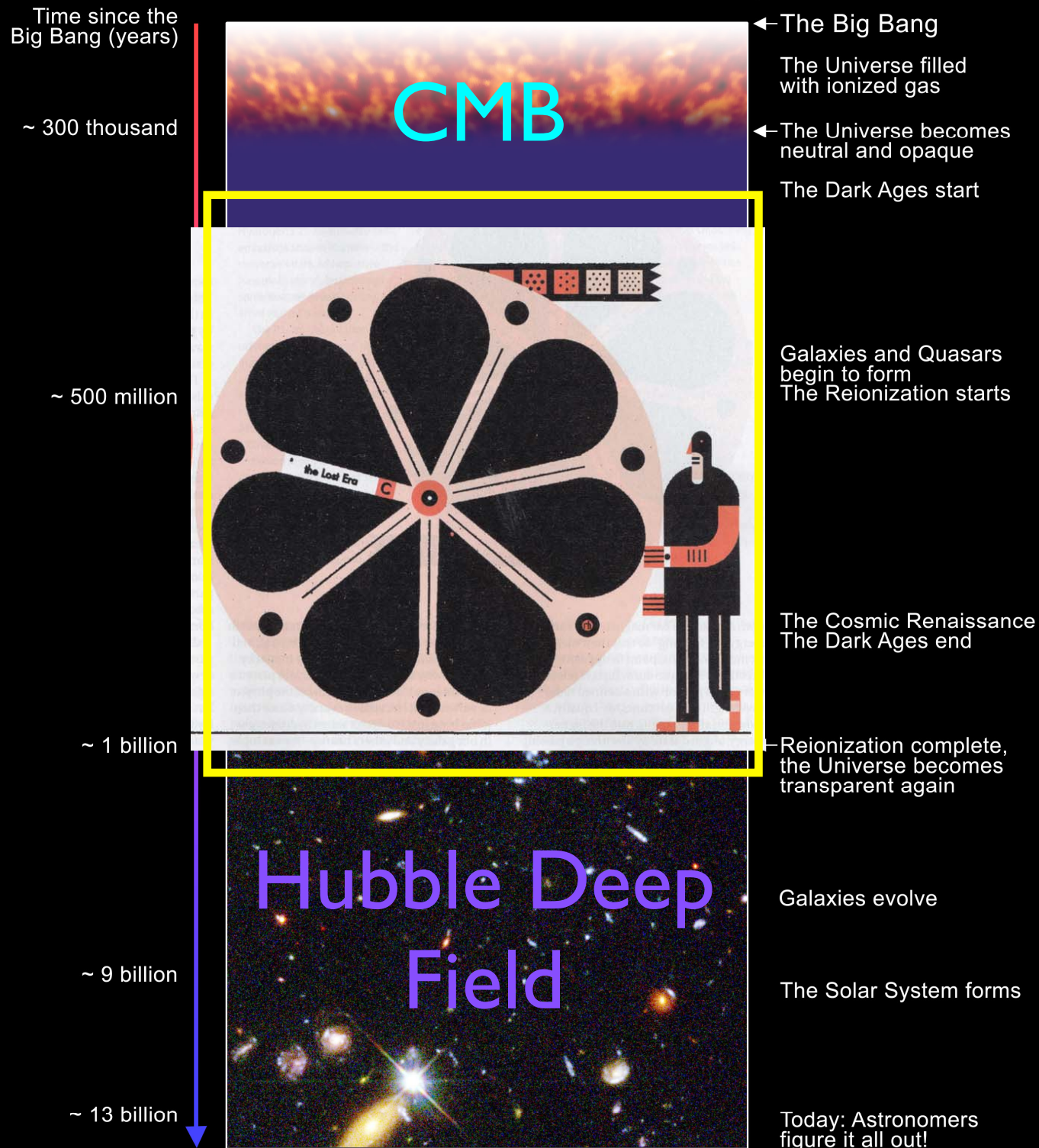
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# The Missing Reel

## What is the Reionization Era?

A Schematic Outline of the Cosmic History

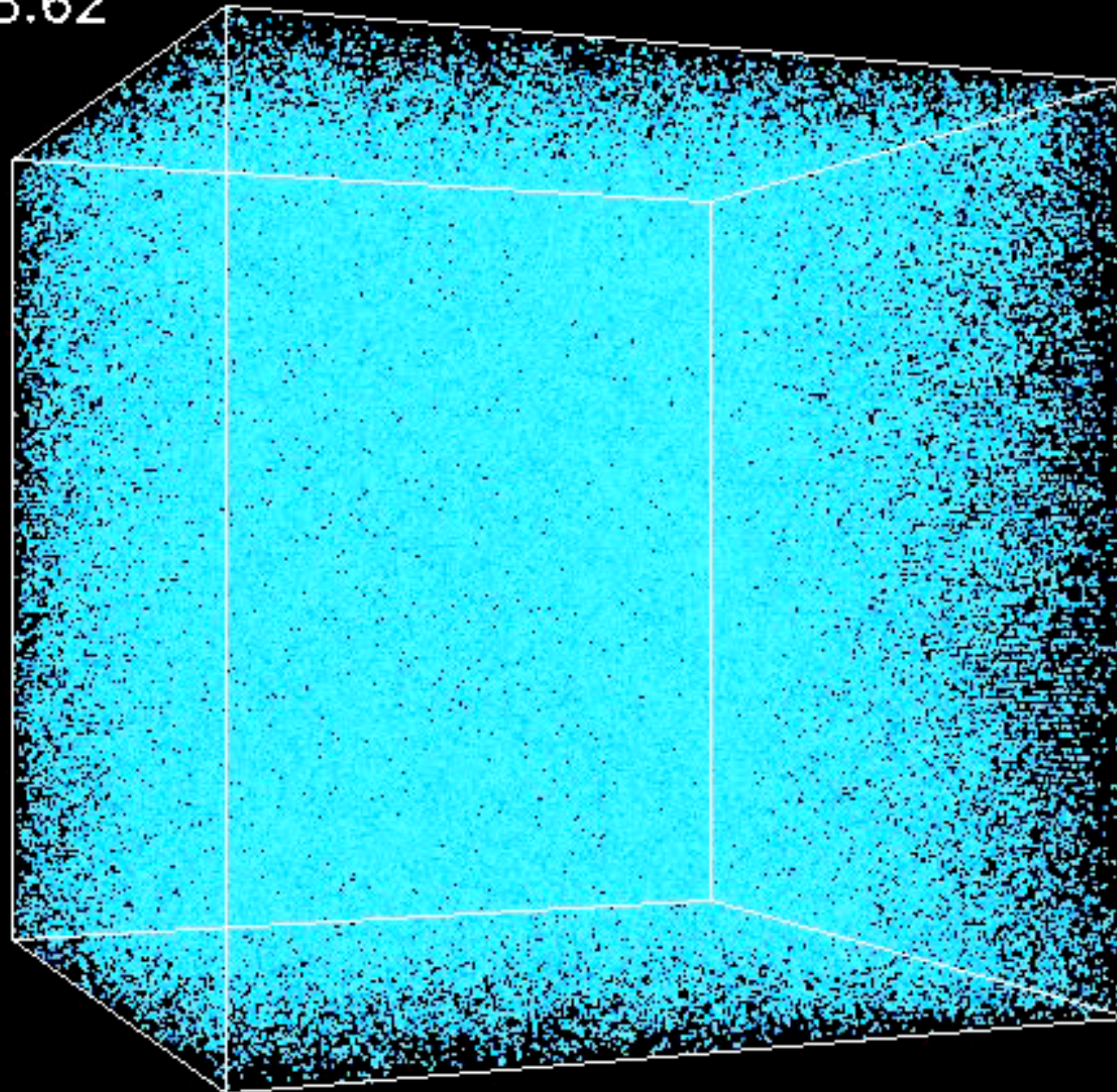


- Dark Ages
- Cosmic Dawn
- Reionization



# Dark sure, but interesting

$Z=28.62$

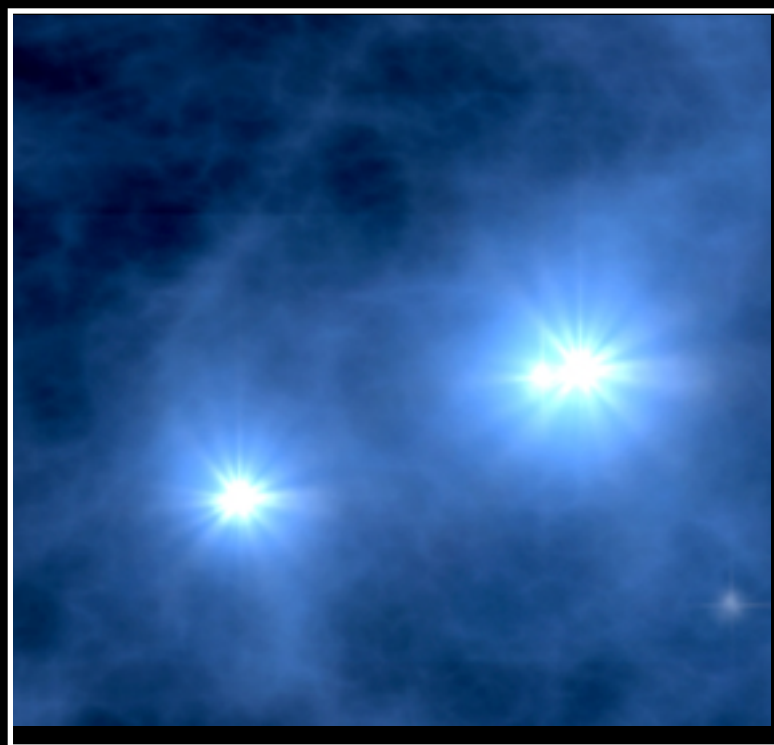


Structures grow as gravity causes collapse



# Fire breathing monsters?

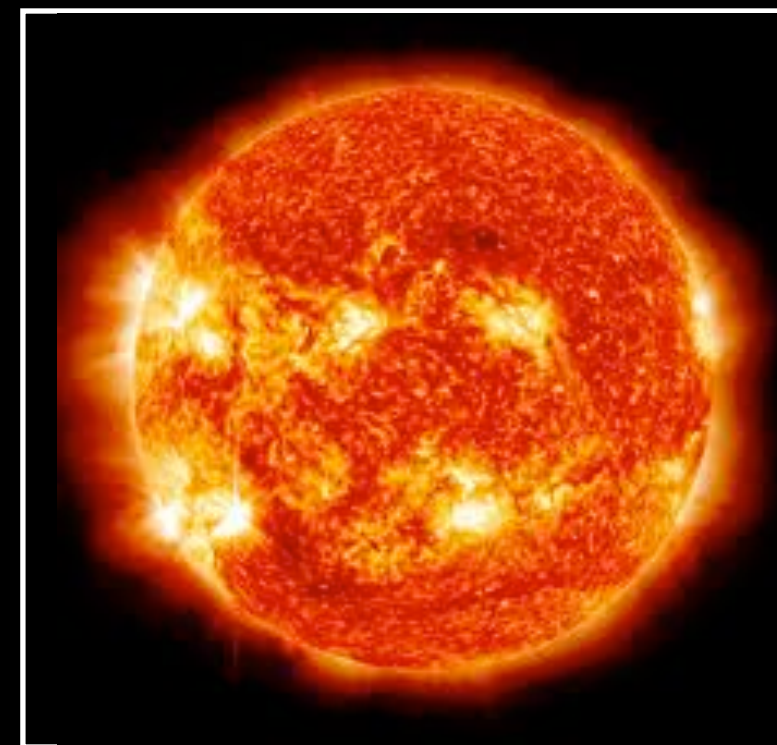
- First stars formed from primordial gas - hydrogen and helium
- Later stars form from gas enriched with metals - carbon, oxygen, iron,...
- First stars may have been larger, hotter, brighter => live fast, die young (Pop III lifetime ~ 1 million years)



Population III  
Metal-free  
~100 Msol  
~20 Lsol



Population II  
Metal-poor  
(galactic bulge/globular clusters)



Population I  
Metal-rich  
(spiral arms of Milky way)  
Our sun



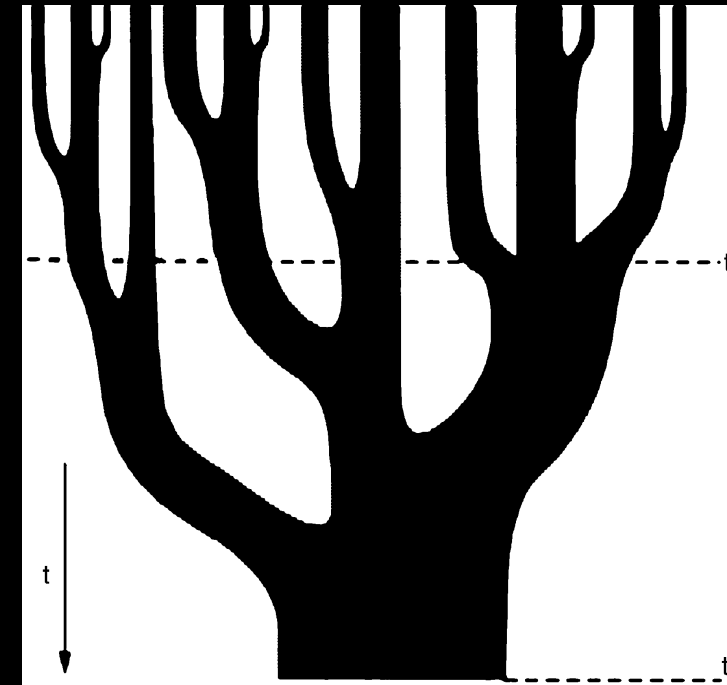
# First stars





# Galactic archaeology

- “hierarchical structure formation”
  - small galaxies merge to form larger ones
- Can search for oldest stars in our galaxy as **low-mass metal-poor stars**
- Less fossils, more very old citizens of Universe

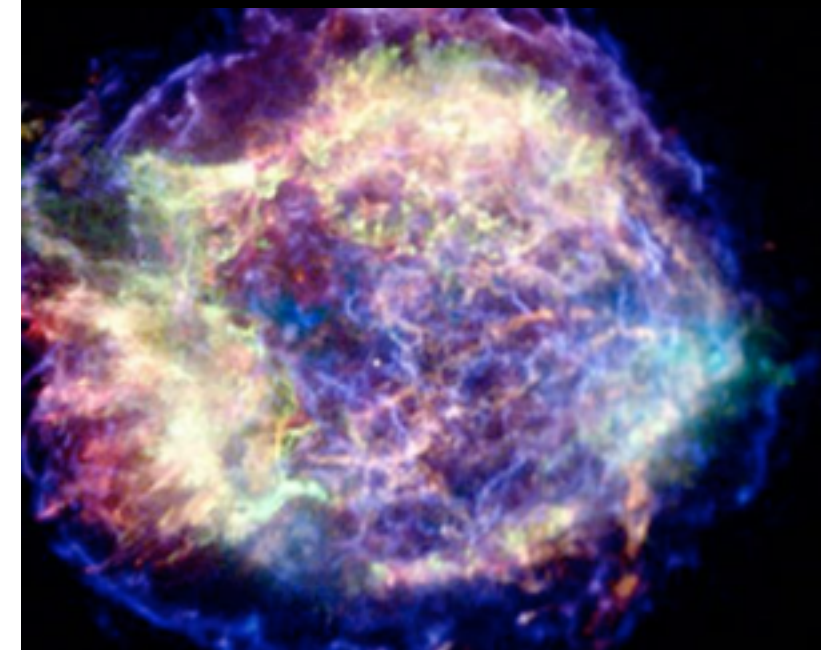


## LETTER

doi:10.1038/nature12990

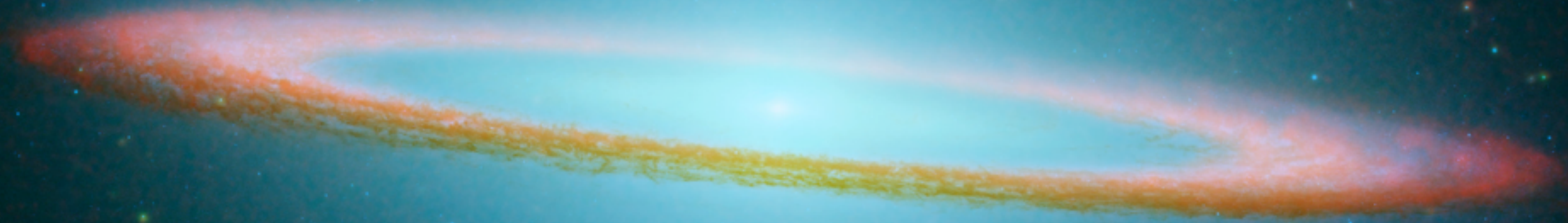
### A single low-energy, iron-poor supernova as the source of metals in the star SMSS J031300.36–670839.3

S. C. Keller<sup>1</sup>, M. S. Bessell<sup>1</sup>, A. Frebel<sup>2</sup>, A. R. Casey<sup>1</sup>, M. Asplund<sup>1</sup>, H. R. Jacobson<sup>2</sup>, K. Lind<sup>3</sup>, J. E. Norris<sup>1</sup>, D. Yong<sup>1</sup>, A. Heger<sup>4</sup>, Z. Magic<sup>1,5</sup>, G. S. Da Costa<sup>1</sup>, B. P. Schmidt<sup>1</sup> & P. Tisserand<sup>1</sup>





# First galaxies

- Galaxy is a collection of gravitationally bound stars
  - First star to form altered surroundings
    - heating, ionization, metal enrichment, winds,...
  - IF feedback prevents star formation for a while then galaxies are “bursty” - 1 star at a time
  - IF feedback mild then may get more continuous star formation
  - Prima dona vs team players...
- 



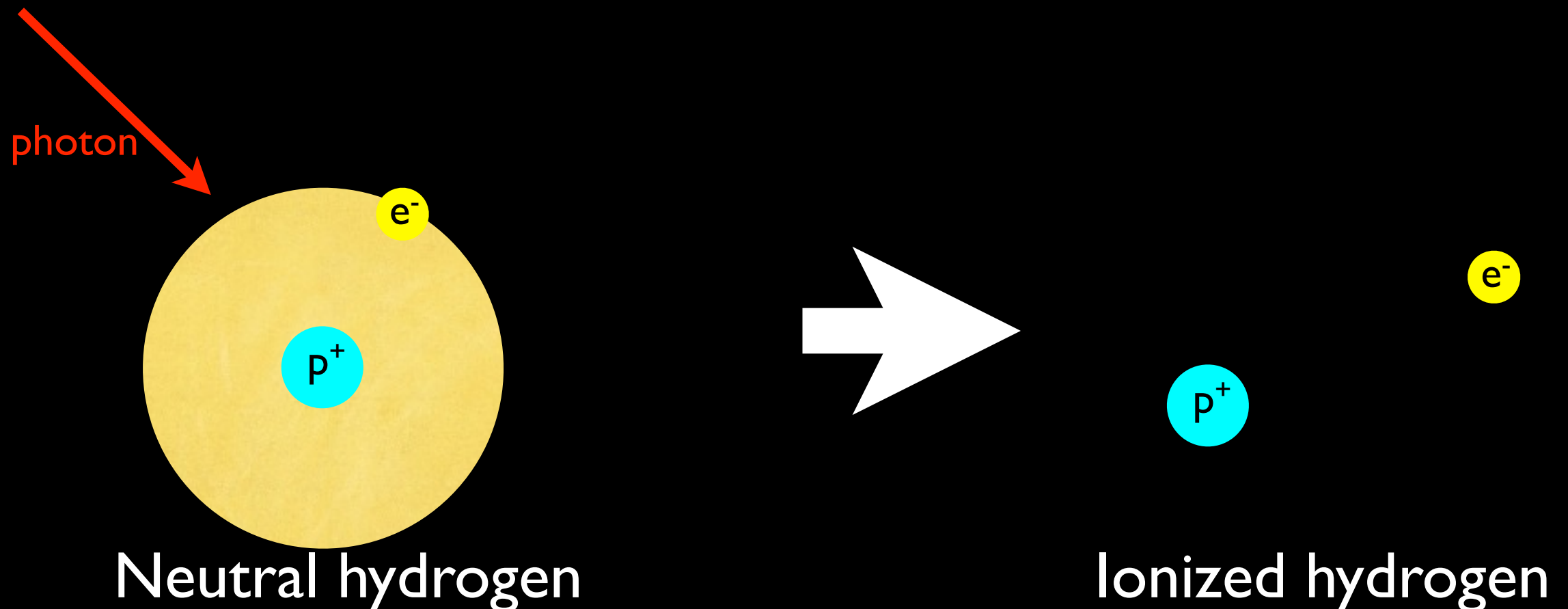
# What sort of light?

- As stars live and die they produce remnants - neutron stars, black holes
- Black holes can coalesce at center of galaxy and grow rapidly => supermassive black hole
- Accretion of gas onto black hole releases energy as radiation - can out shine galaxy = Active Galactic Nuclei (AGN)
- Jet of non-thermal emission + hot X-rays



# Blowing bubbles

- Light more energetic than 13.6 eV can ionize hydrogen
- Ionized bubbles form around galaxies and grow, merging over time.
- Ultimately, the space between galaxies becomes entirely ionized





# Reionization



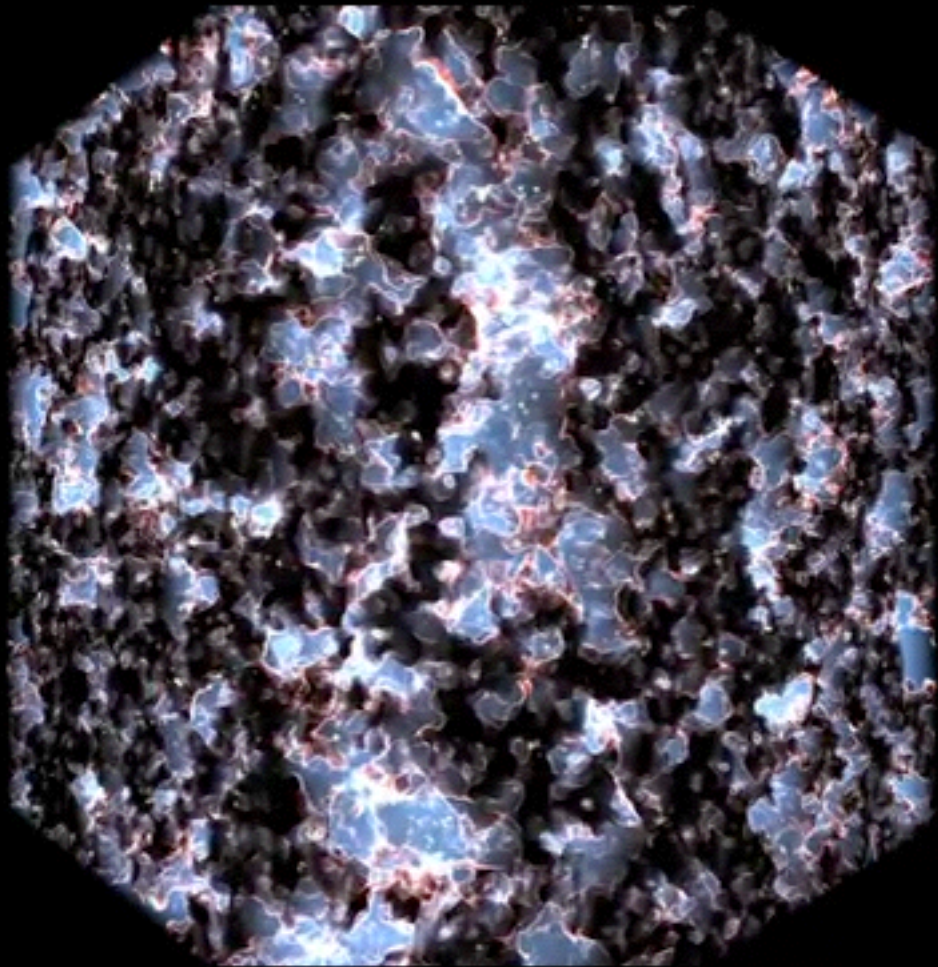


# The last phase transition???

- Reionization is the last major phase transition in the Universe from space being filled with cold, neutral gas to hot, ionized gas.
- When and how long did reionization take place?
- What were the sources that drove it?
  - Massive metal-free Population III stars
  - Many metal-poor Population II stars
  - Accretion onto supermassive black holes (AGN)
- How did ionized bubbles grow and merge?
  - small bubbles around individual galaxies
  - larger bubbles around groups of galaxies



# What do we know?



Reionization





# Reionization

Cast:

dark matter halos

Population III stars

Population II stars

Galaxies

Black holes

Dark stars

Miniquasar

...

# Casablanca

Cast:

Taxi driver

Police man

Rick

Souk merchant

Lazlo

Ilsa

Man with camel

...

**Know the cast, but  
who are the leads and who the bit players?**



# Reionization

Plot:

Universe was neutral

Luminous sources form

Universe became ionized

# Casablanca

Plot:

Boy meets girl

Boy falls for girl

Boy loses girl

Know the plot highlights,  
but what are the details?





# Cosmic movie making

---

**DIRECTOR** T. Universe

---

**CAMERAMAN** S.K.A.

---

<b>DATE</b>	<b>TAKE</b>	<b>SCENE</b>
13.75 GYrABB	1	$10^{51}$

---

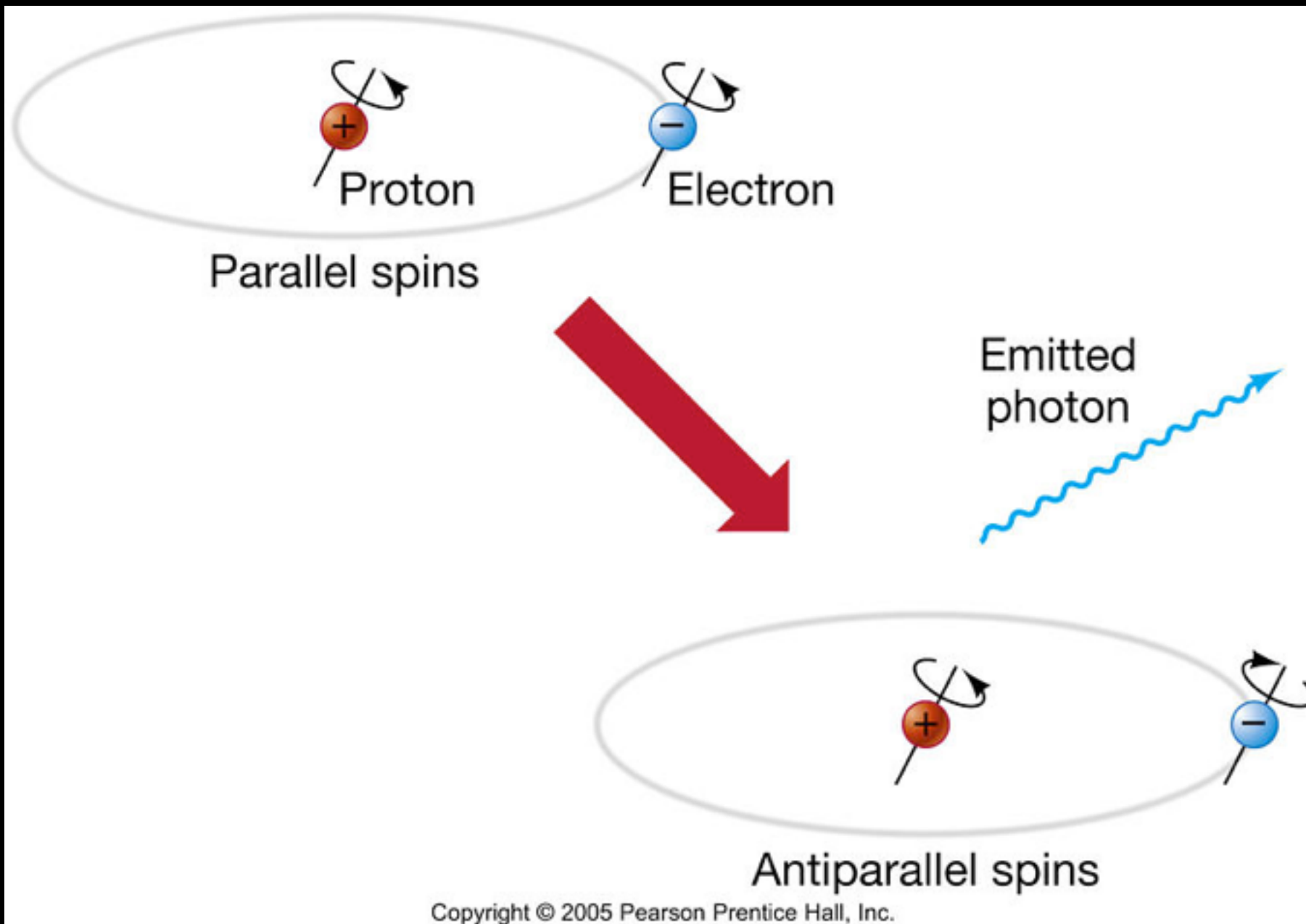


# How to observe reionization?

- Should be able to see some of the bright galaxies with the James Webb Space Telescope  
**BUT**  
faint galaxies are too faint. Sad since they're numerous
- Really want to see the hydrogen gas that gets ionized  
**BUT**  
light from ionization and recombination of hydrogen gets blocked before it gets to us
- Need a new technique to map the hydrogen

# The 21 cm line

- Hydrogen can emit or absorb radio waves with wavelength of 21cm from spin flips
- These radio waves can propagate to us today

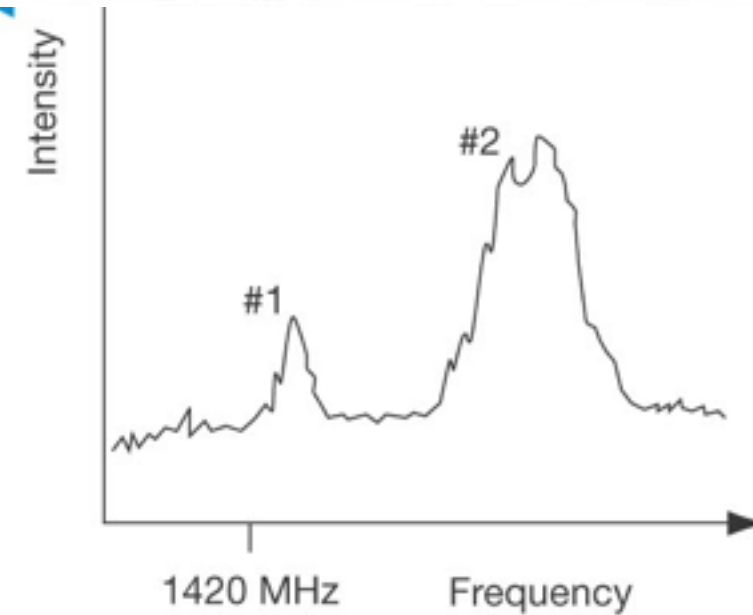
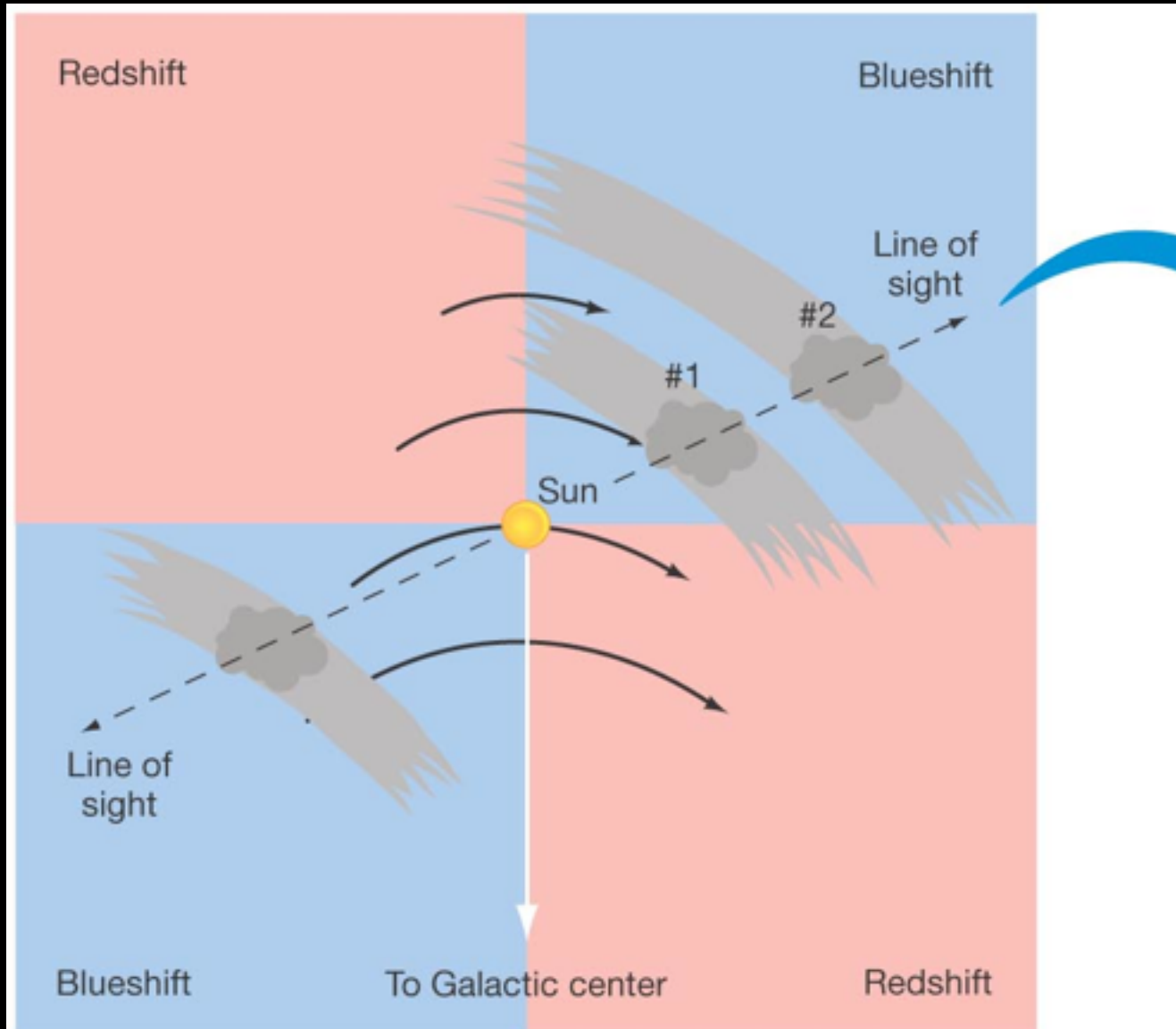
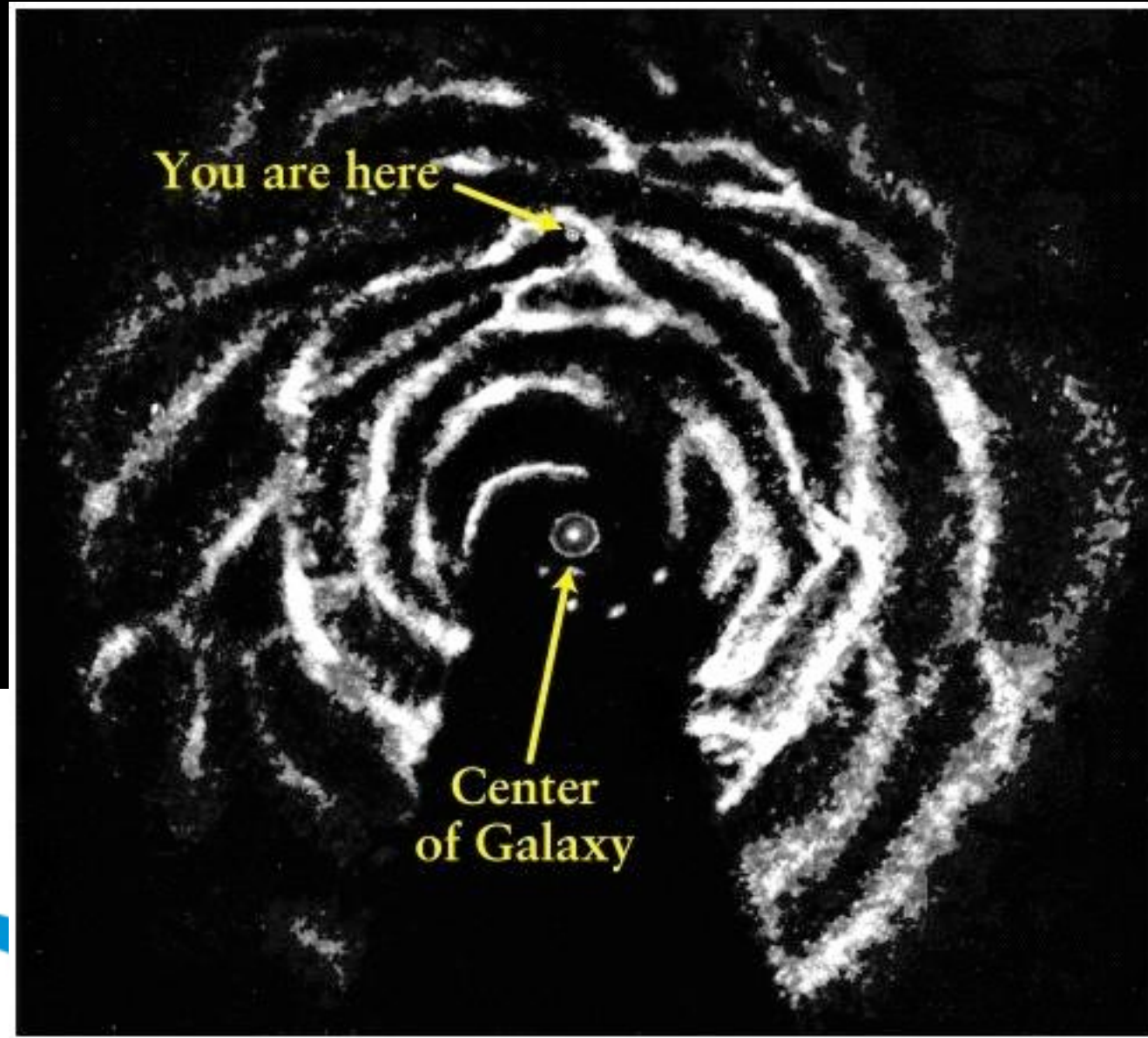


21cm = 1.4 GHz



# Nearby 21 cm

- 21 cm light will be shifted in
- Used in nearby Universe to

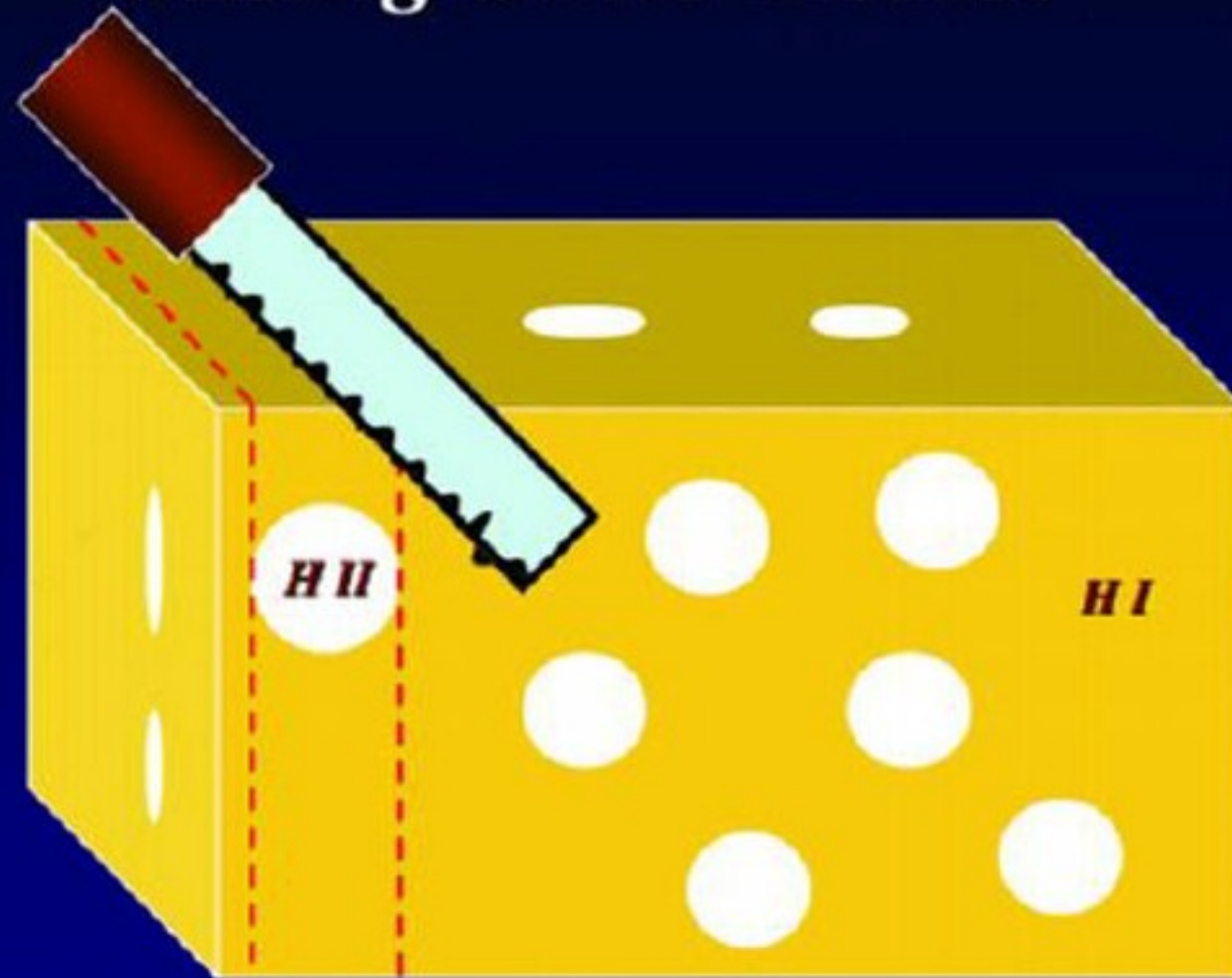


# Cutting Swiss Cheese

- At cosmic distances, expansion of Universe produces Doppler shift = redshifting

*21cm Tomography of Ionized Bubbles During Reionization is like*

*Slicing Swiss Cheese*



*Observed wavelength  $\leftrightarrow$  distance*

$$21\text{cm} \times (1 + z)$$



# Where should we look?

- Frequencies **40-200MHz** correspond to **100million - 1 billion years** after the Big Bang

$$1420 \text{ MHz} \Rightarrow z=0$$

$$t_{\text{Age}}(z=0) \sim 13.7 \text{ Gyr}$$

$$200 \text{ MHz} \rightarrow z = 6$$

$$t_{\text{Age}}(z = 6) \approx 1 \text{ Gyr}$$

$$100 \text{ MHz} \rightarrow z = 13$$

$$t_{\text{Age}}(z = 10) \approx 500 \text{ Myr}$$

$$70 \text{ MHz} \rightarrow z \approx 20$$

$$t_{\text{Age}}(z = 20) \approx 150 \text{ Myr}$$

$$50 \text{ MHz} \Rightarrow z \sim 27$$

$$t_{\text{Age}}(z=27) \sim 100 \text{ Myr}$$

# Radio telescopes of the past

## NEW RADIO WAVES TRACED TO CENTRE OF THE MILKY WAY

Mysterious Static, Reported  
by K. G. Jansky, Held to  
Differ From Cosmic Ray.

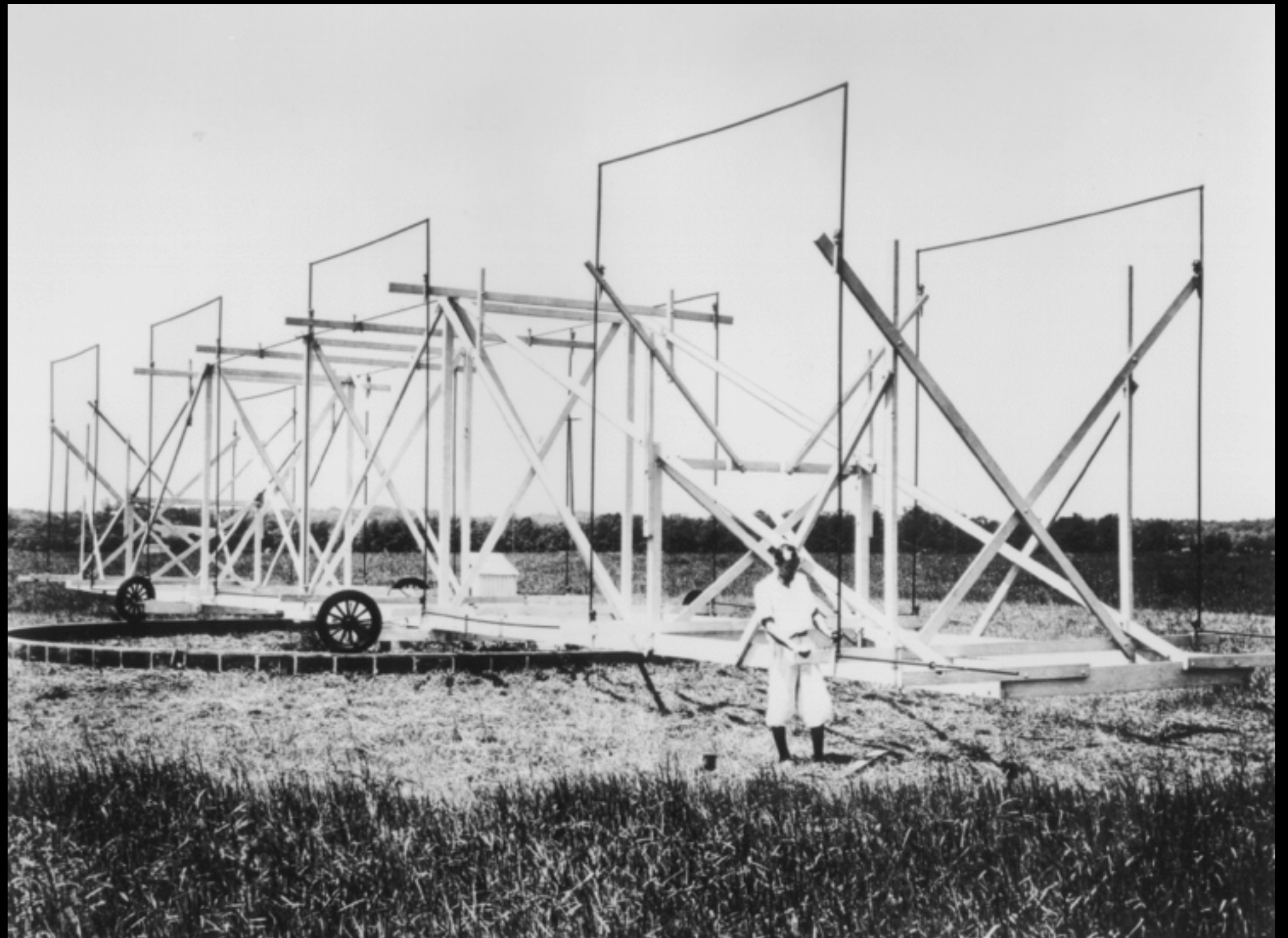
DIRECTION IS UNCHANGING

Recorded and Tested for More  
Than Year to Identify it as  
From Earth's Galaxy.

ITS INTENSITY IS LOW

Only Delicate Receiver is Able to  
Register—No Evidence of  
Interstellar Signaling.

Discovery of mysterious radio waves which appear to come from the centre of the Milky Way galaxy was announced yesterday by the Bell Telephone Laboratories. The discovery was made during research studies on static by Karl G. Jansky of the radio research department at Holmdel, N. J., and was described by him in a paper delivered before the International Scientific Radio Union in Washington.



Front page  
New York Times  
May 3, 1933

Jansky with his 20.5 MHz telescope near Bell Labs



# Bigger is better!



Green Bank 300ft telescope



# Building big dishes is hard

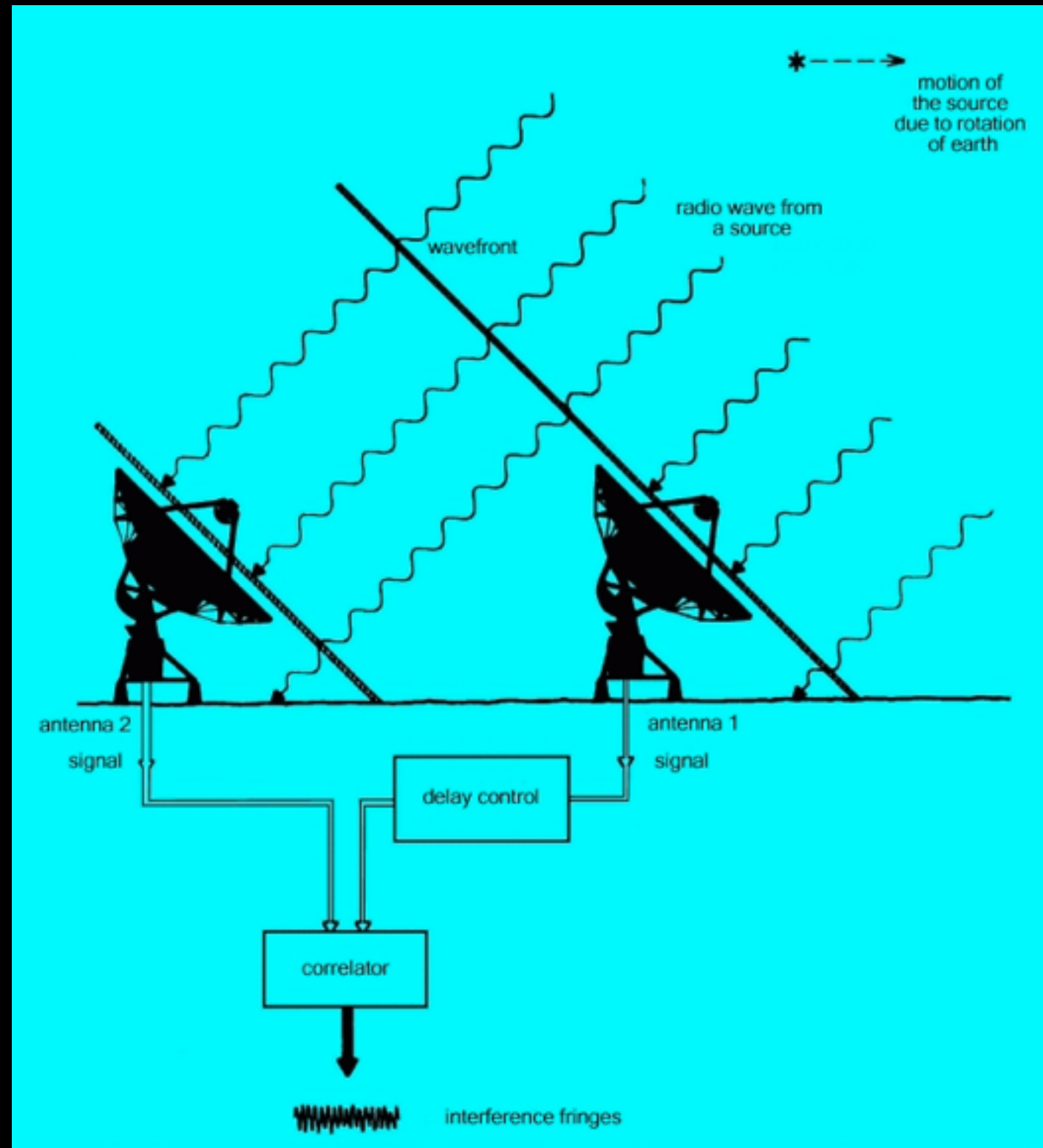


collapsed in 1988 - structural failure



# Interferometry

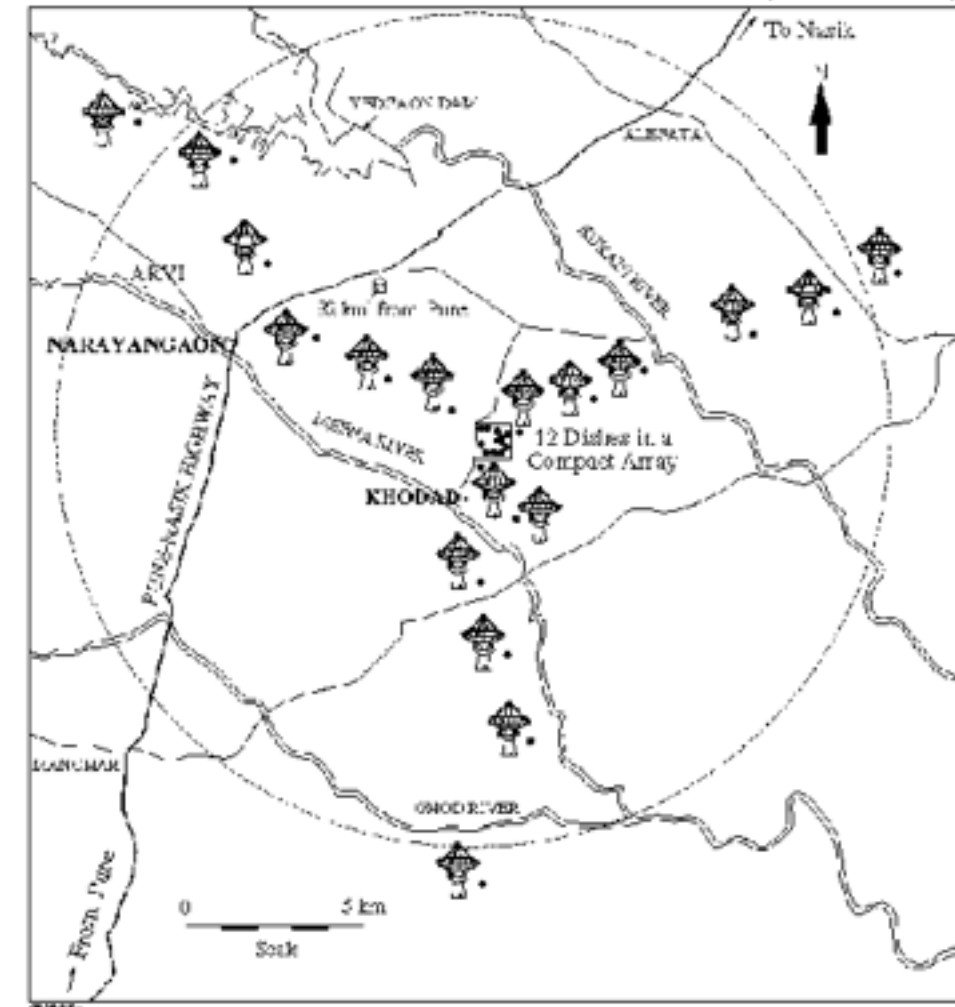
- Radio telescope samples phase and amplitude of electromagnetic wave
- Combine signal from many elements to synthesise a bigger telescope = 'interferometry'
- More elements = more collecting area & better quality image



# Giant Meter Radio Telescope



LOCATIONS OF GMRT ANTENNAS ( 30 dishes )

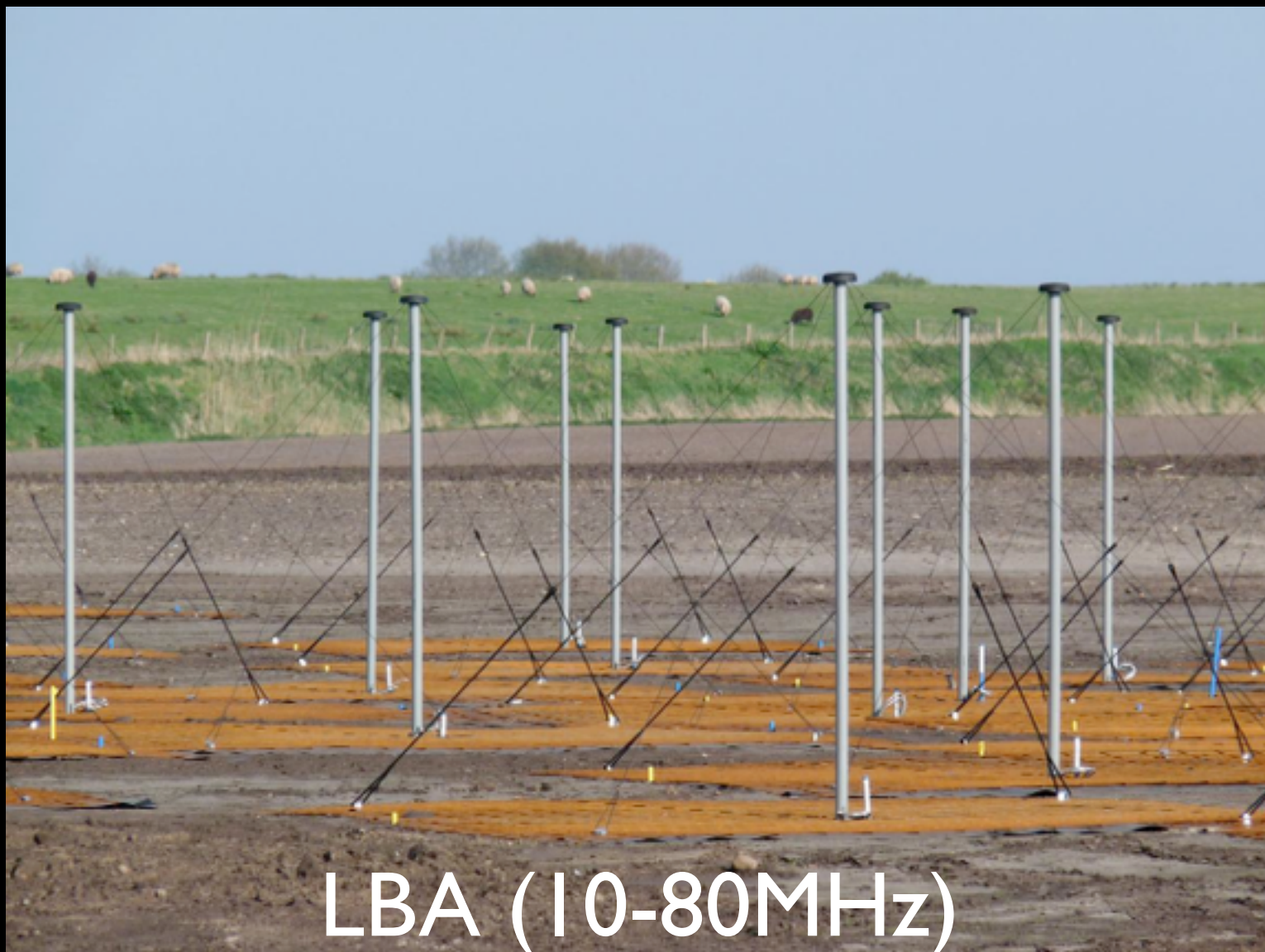


Built in 1995, 80 km from Pune, India  
30 x 45m dishes. 50MHz-1500MHz



# LOW Frequency ARray (LOFAR)

- First of modern arrays built to target cosmic 21 cm signal
- Individual elements are simple antennae collected into 'tiles'
- Plug into supercomputer and correlate = 'Digital astronomy'



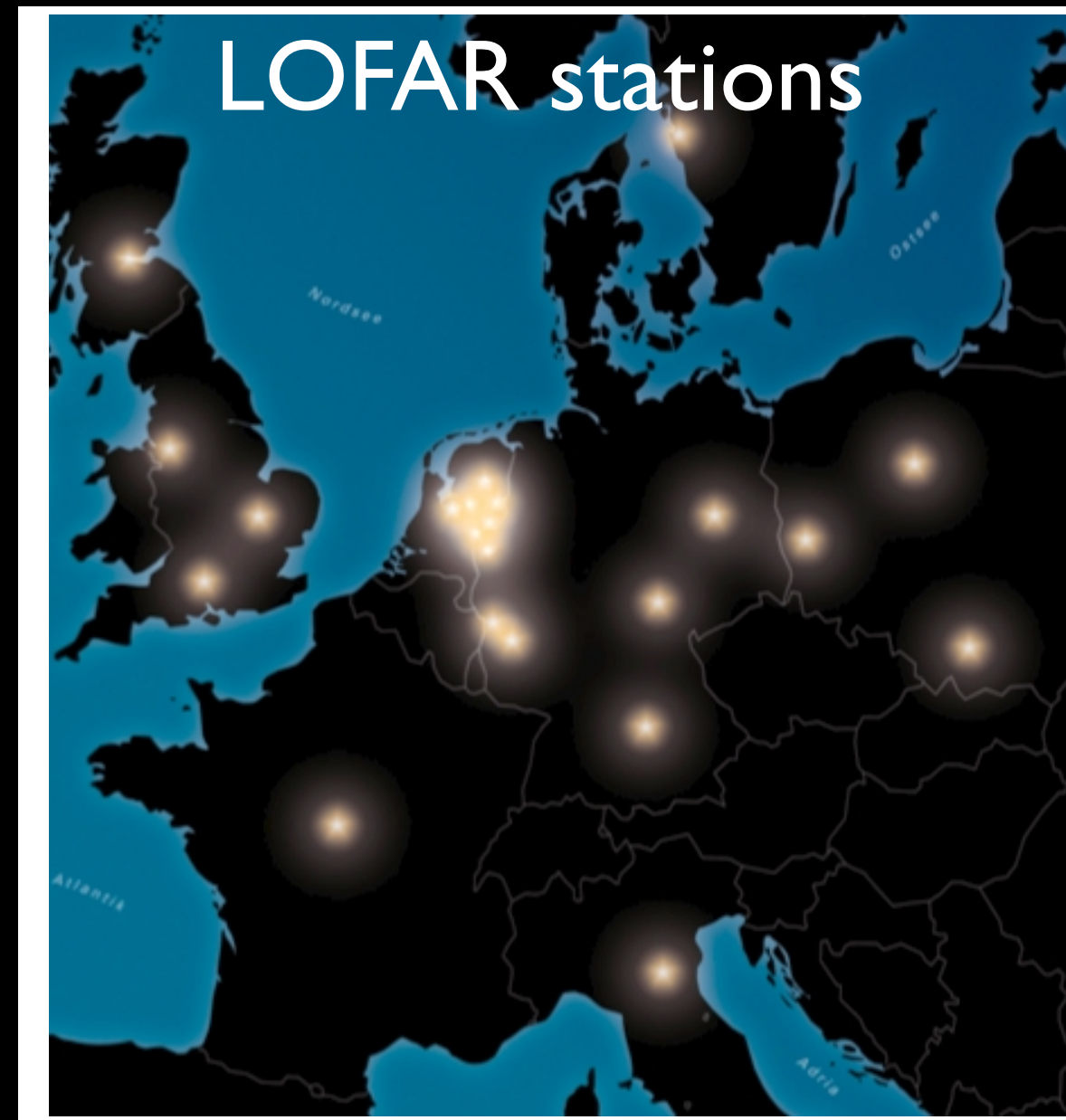
## HBA (120-240 MHz)





# LOFAR

- Core located in Netherlands
- Stations scattered over Europe
- Now looking for cosmic 21 cm signal





# Square Kilometer Array

- Will have ~ 1 million antennae and 1km<sup>2</sup> collecting area
- Distributed over core of ~ 1km to outlying stations ~ 100km away
- Located in South Africa and Australia
- First light ~2020
- International HQ at Jodrell Bank near Manchester
- Cost: €650 million for Phase I
- BIG DATA!!! - SKA Phase I will produce today's internet per night



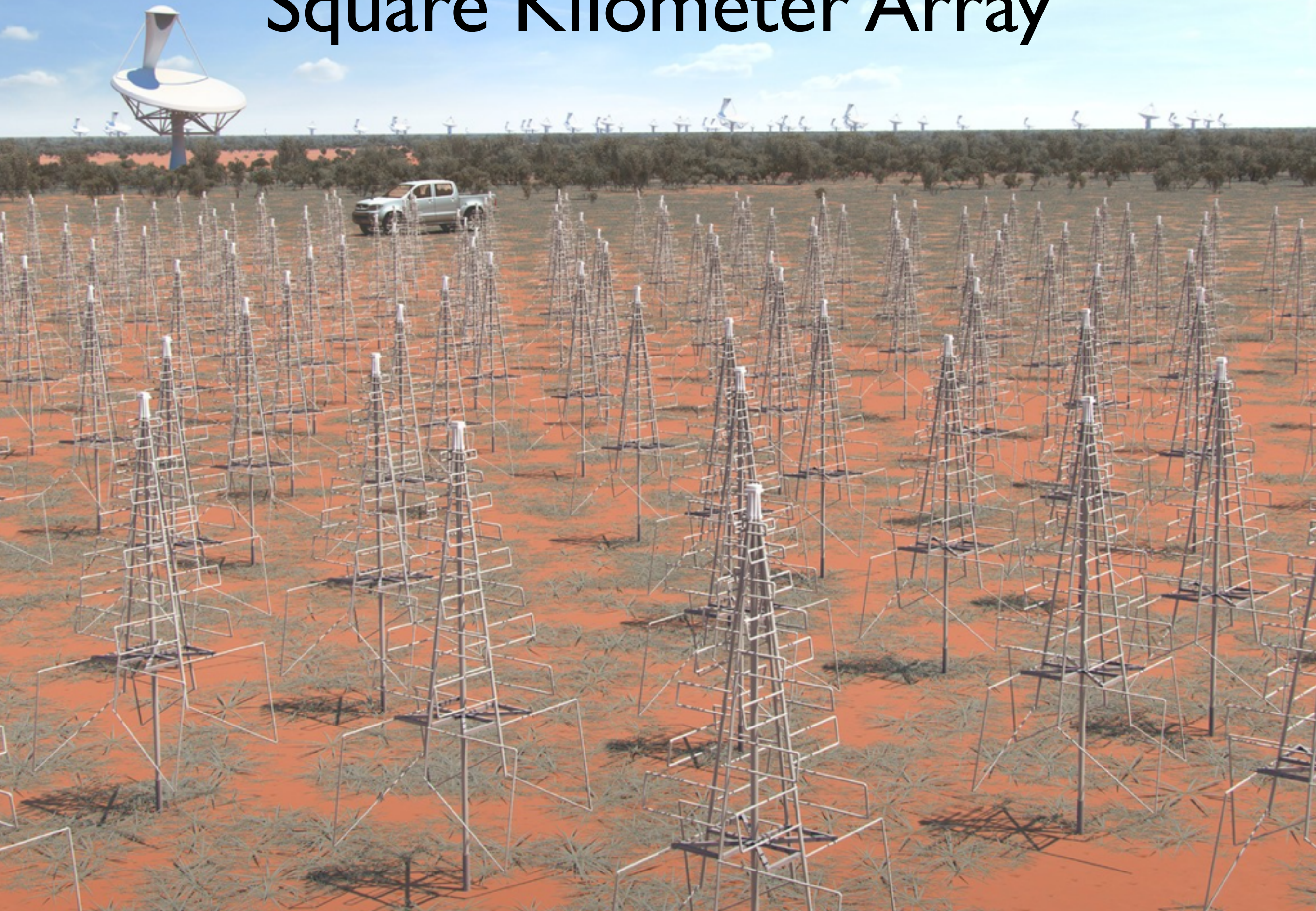
# Square Kilometer Array



THE SQUARE KILOMETRE ARRAY



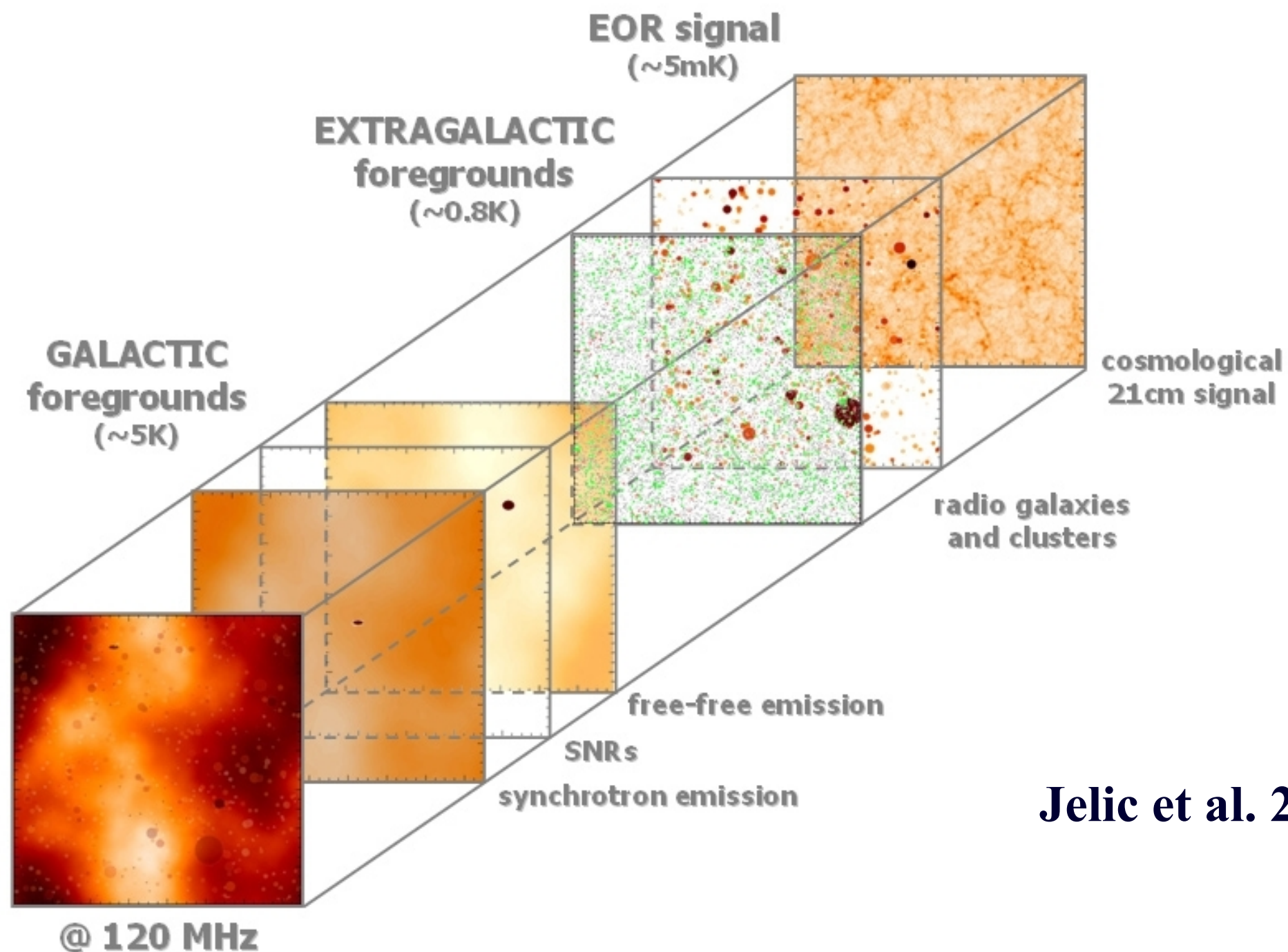
# Square Kilometer Array





# Key challenges

- Man made radio interference
- Astronomical foregrounds - 1000x larger than signal



Jelic et al. 2008

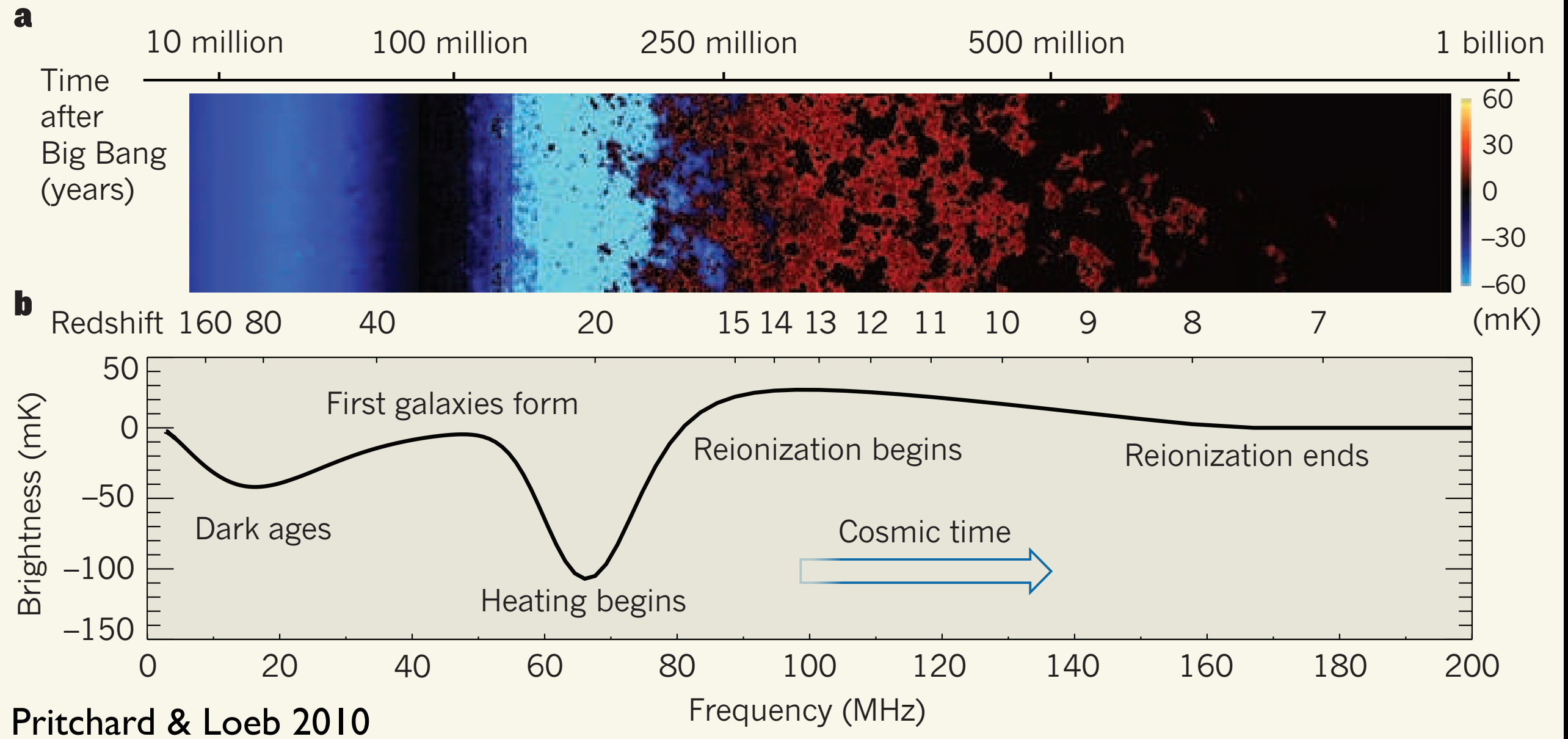


# 21 cm signal

Moon?

SKA

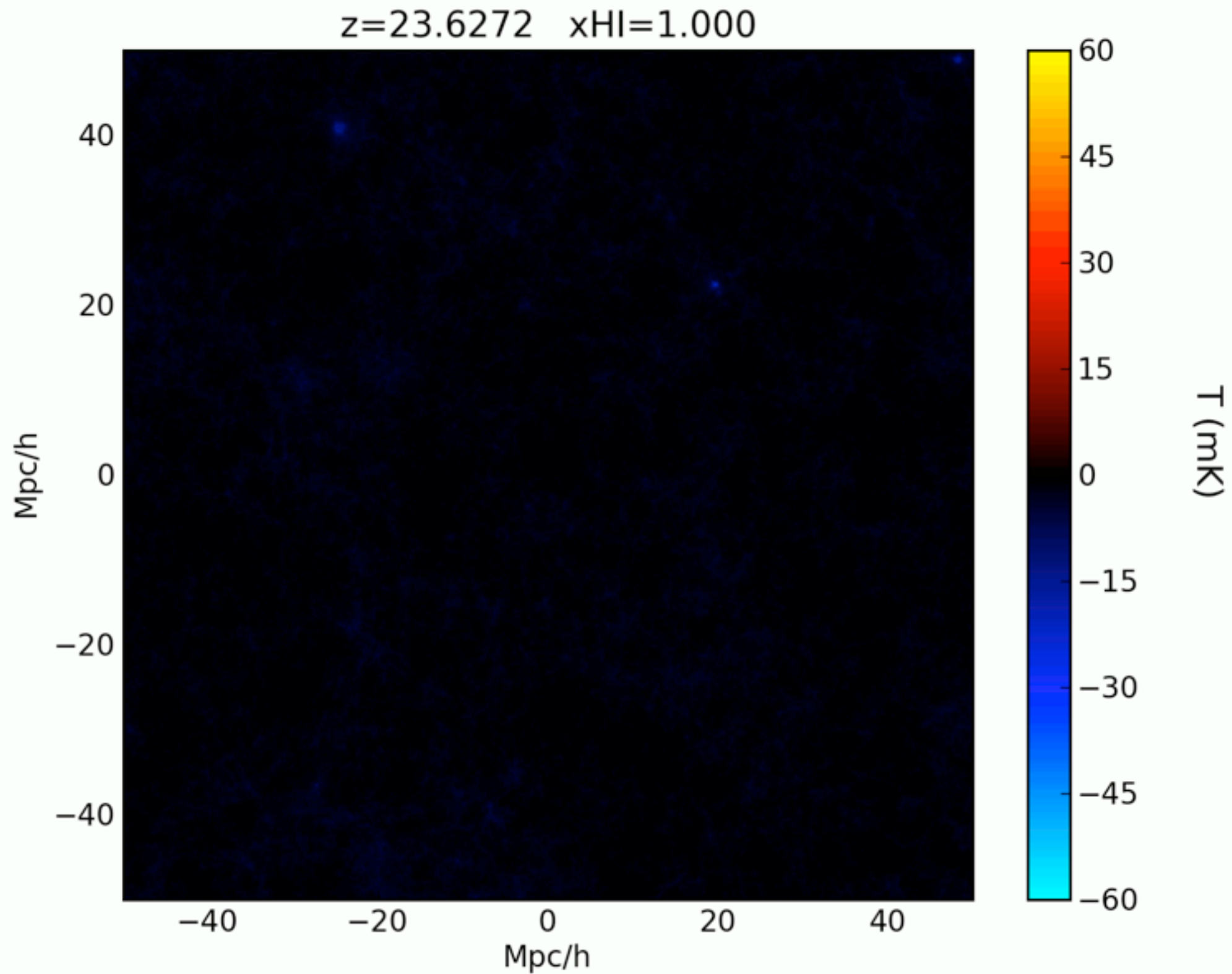
LOFAR/MWA/PAPER



Where CMB is a photo; 21 cm signal is a movie

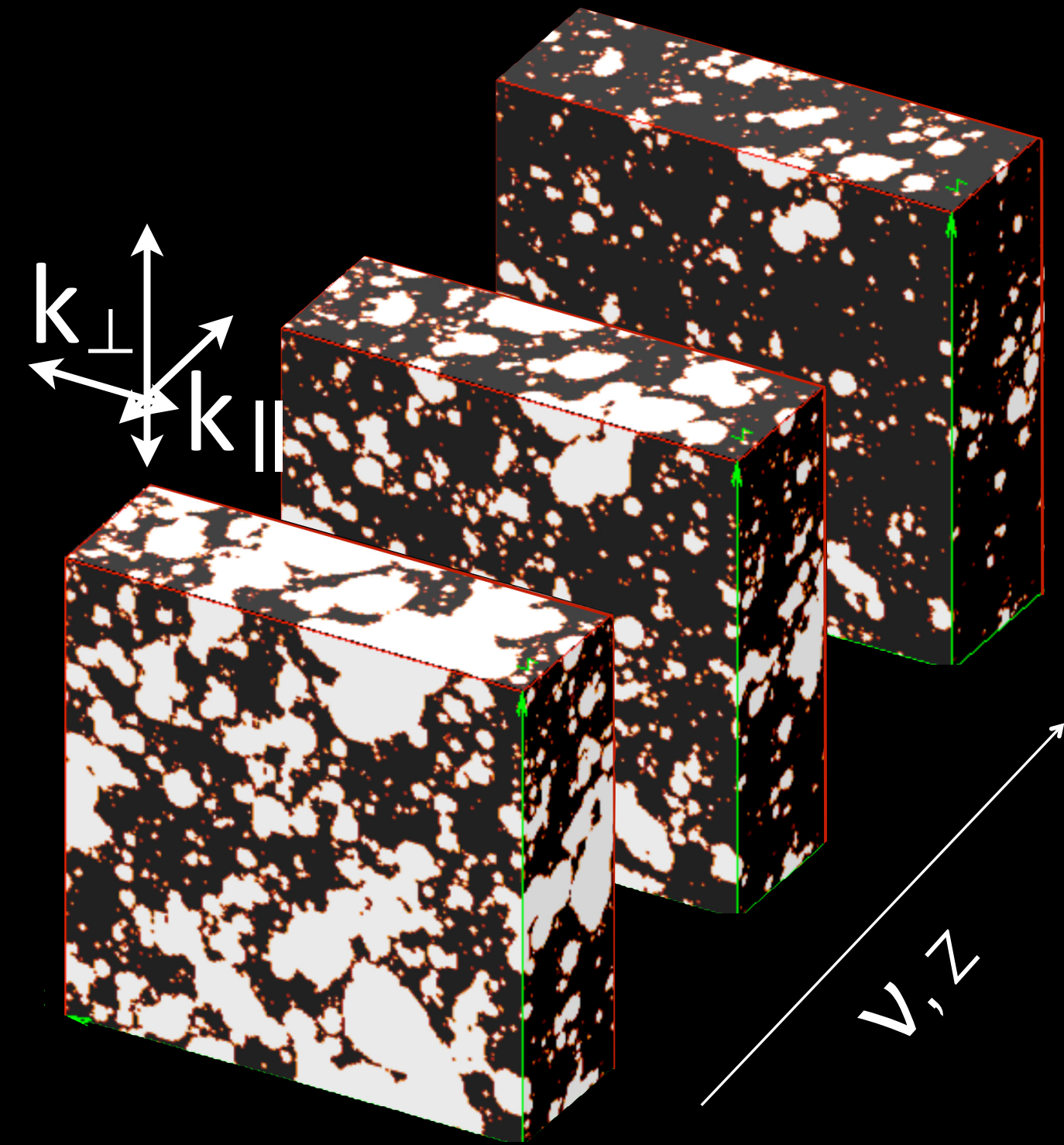
Sensitive to key moments in cosmic history

# Cosmic movies



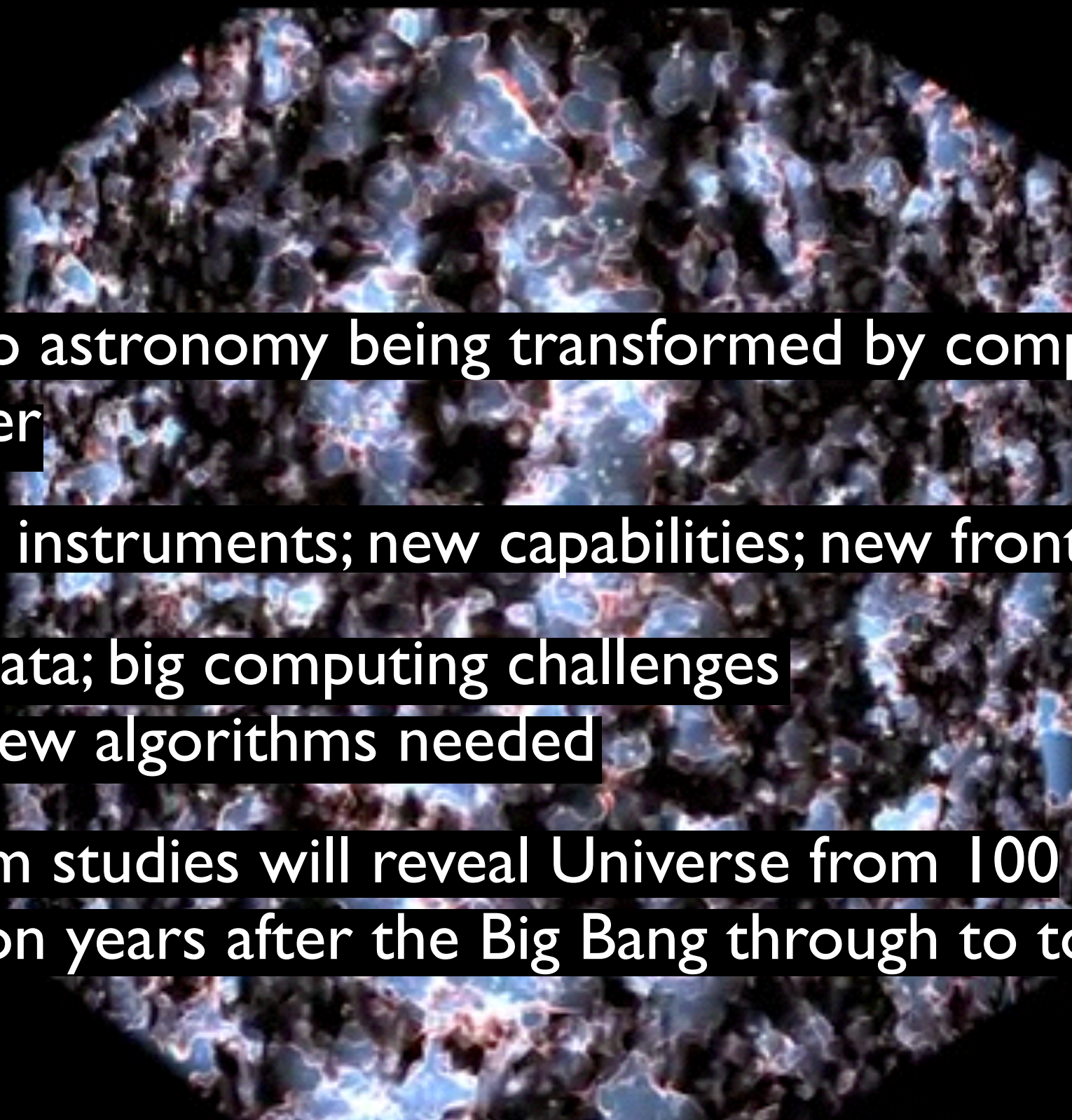


# What might we learn?



- When did first galaxies begin heating and ionizing the Universe.
- Sizes & number of bubbles linked to brightness and abundance of galaxies
- Temperature of hydrogen linked to X-ray emission from early black holes
- Map distribution of matter throughout Universe

# Things to watch out for

- 
- Radio astronomy being transformed by computing power
  - New instruments; new capabilities; new frontiers
  - Big data; big computing challenges  
=> new algorithms needed
  - 21 cm studies will reveal Universe from 100 million years after the Big Bang through to today

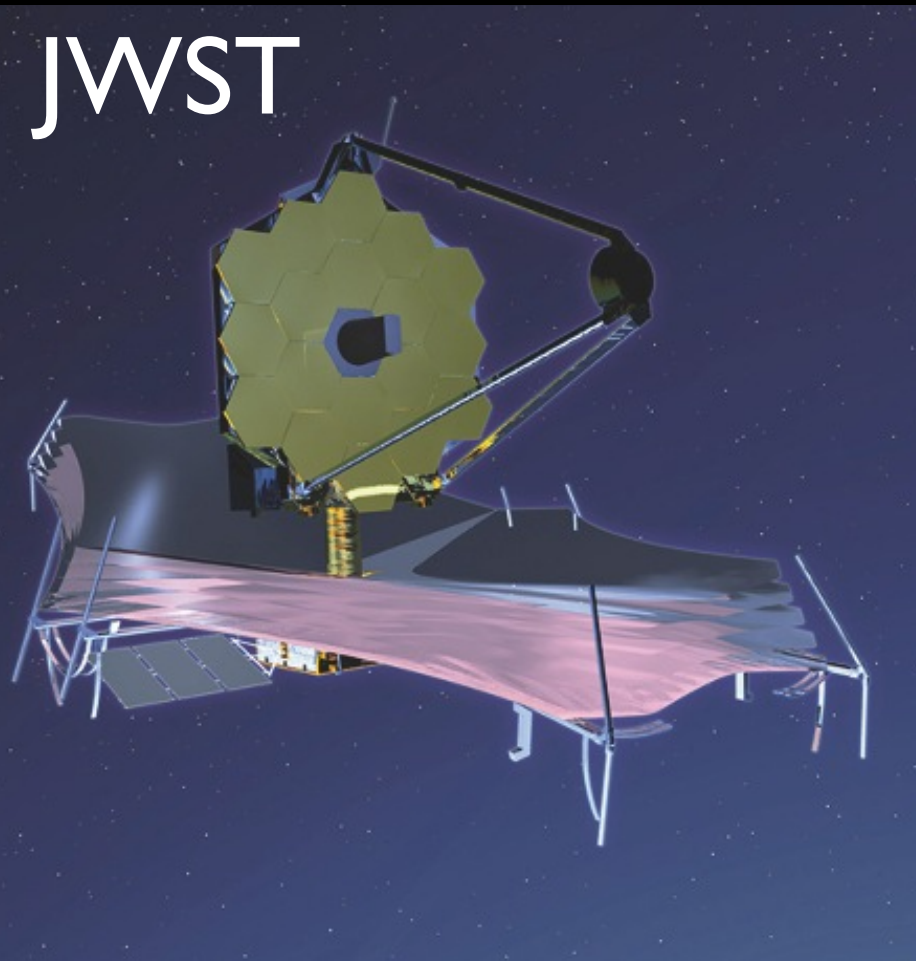


# Exciting times ahead!

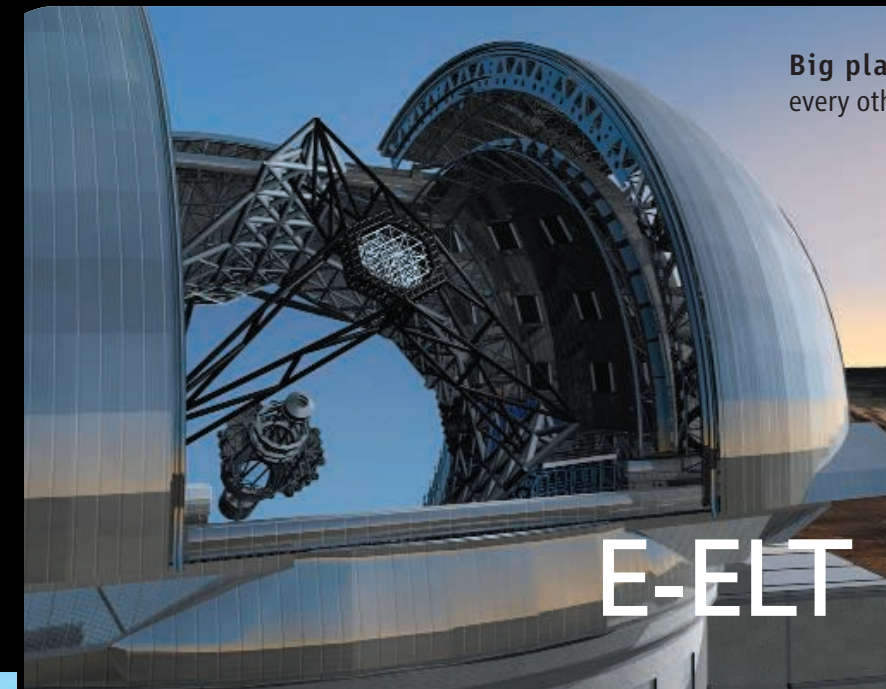
ALMA



JWST



Big plan  
every oth



E-ELT



TMT

GMT





## We are listening - Diane Ackerman

As our metal eyes wake  
to absolute night,  
where whispers fly  
from the beginning of time,  
we cup our ears to the heavens.  
We are listening

on the volcanic rim of Flagstaff  
and in the fields beyond Boston,  
in a great array that blooms  
like coral from the desert floor,  
on highwire webs patrolled  
by computer spiders in Puerto Rico.

We are listening for a sound  
beyond us, beyond sound,

searching for a lighthouse  
in the breakwaters of our uncertainty,  
an electronic murmur,  
a bright, fragile *I am*.

Small as tree frogs  
staking out one end  
of an endless swamp,  
we are listening  
through the longest night  
we imagine, which dawns  
between the life and times of stars.

Our voice trembles  
with its own electric,  
we who mood like iguanas,  
we who breathe sleep  
for a third of our lives,  
we who heat food  
to the steaminess of fresh prey,  
then feast with such  
good manners it grows cold.

In mind gardens  
and on real verandas  
we are listening,  
rapt among the Persian lilacs  
and the crickets,  
while radio telescopes  
roll their heads, as if in anguish.

With our scurrying minds  
and our lidless will  
and our lank, floppy bodies  
and our galloping yens  
and our deep, cosmic loneliness  
and our starboard hearts  
where love careens,  
we are listening,  
the small bipeds  
with the giant dreams.





In the beginning of the Dark Ages, electrically neutral hydrogen gas filled the universe. As stars formed, they ionized the regions immediately around them, creating bubbles here and there. Eventually these bubbles merged together, and intergalactic gas became entirely ionized.

# Thank you for your consideration









# SKA status



SKA director general - Phil Diamond

SKA sites: South Africa - Karoo

Australia - Western Outback

SKA-low: 50-350 MHz (Australia)

SKA-mid: 350MHz-3GHz (SA)

Baseline design currently under discussion

## Current



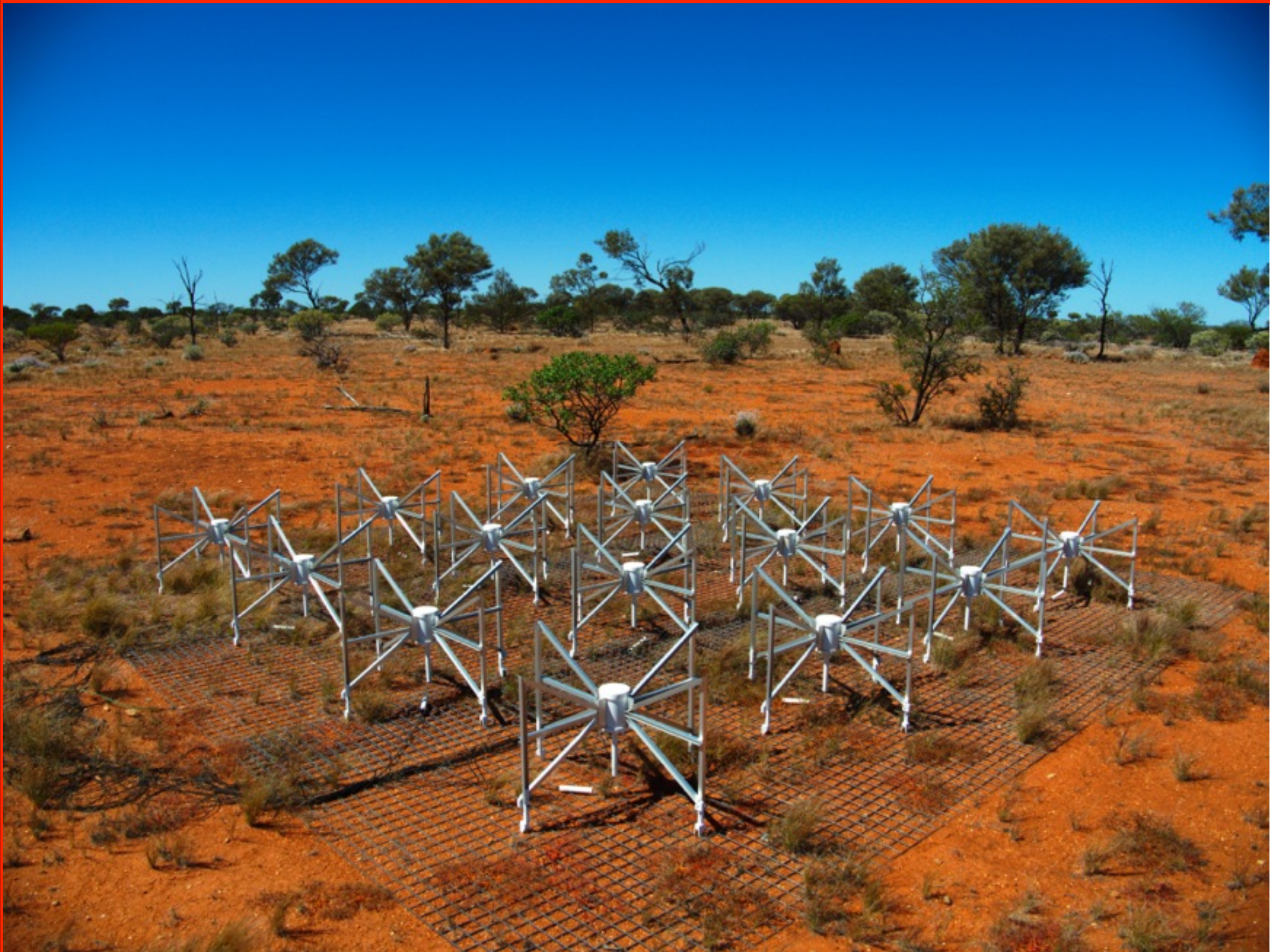
SKA project office - Jodrell Bank, Manchester







# MWA





# Arecibo



305m dish



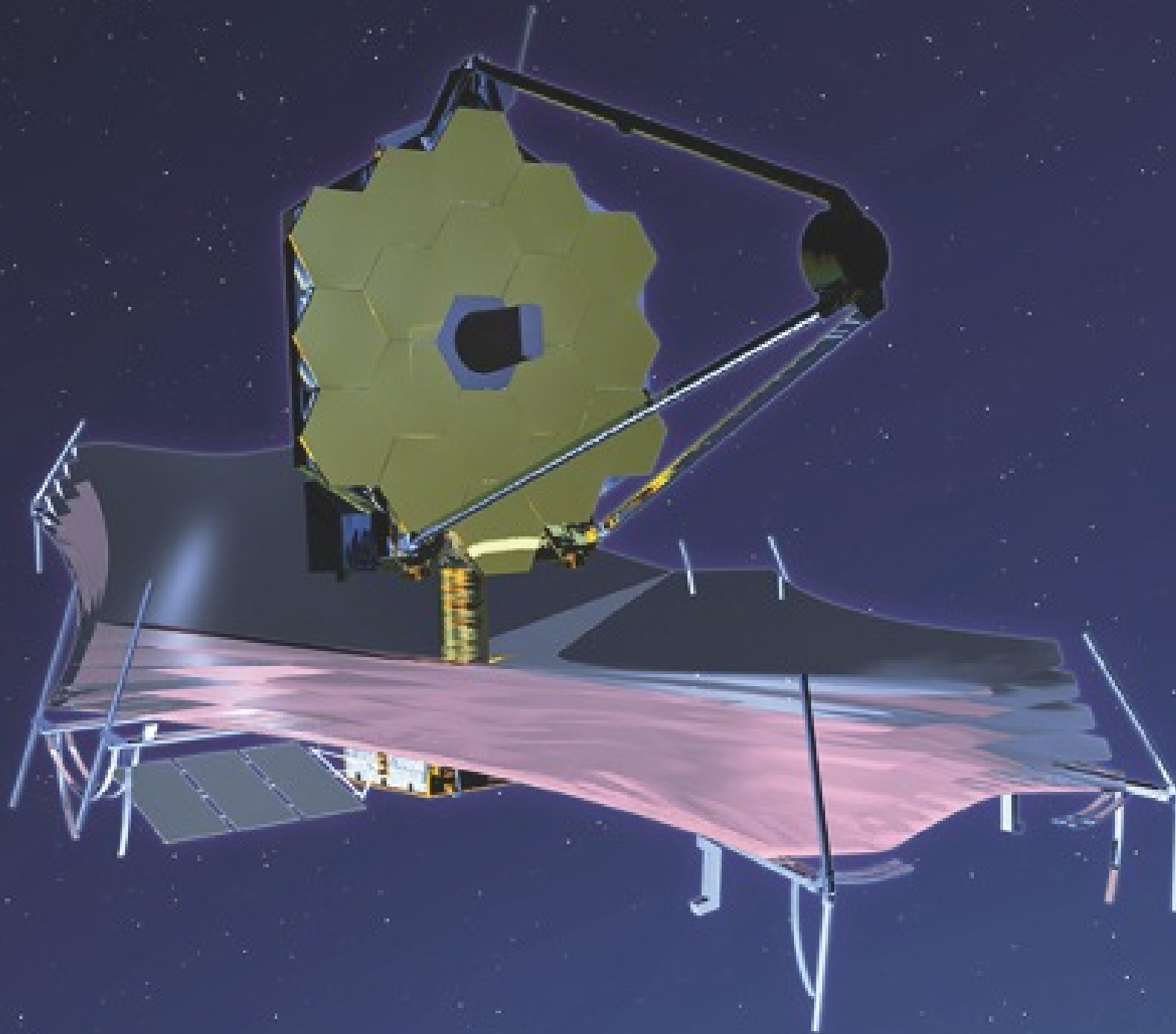


# VLA





# JWST







# ALMA





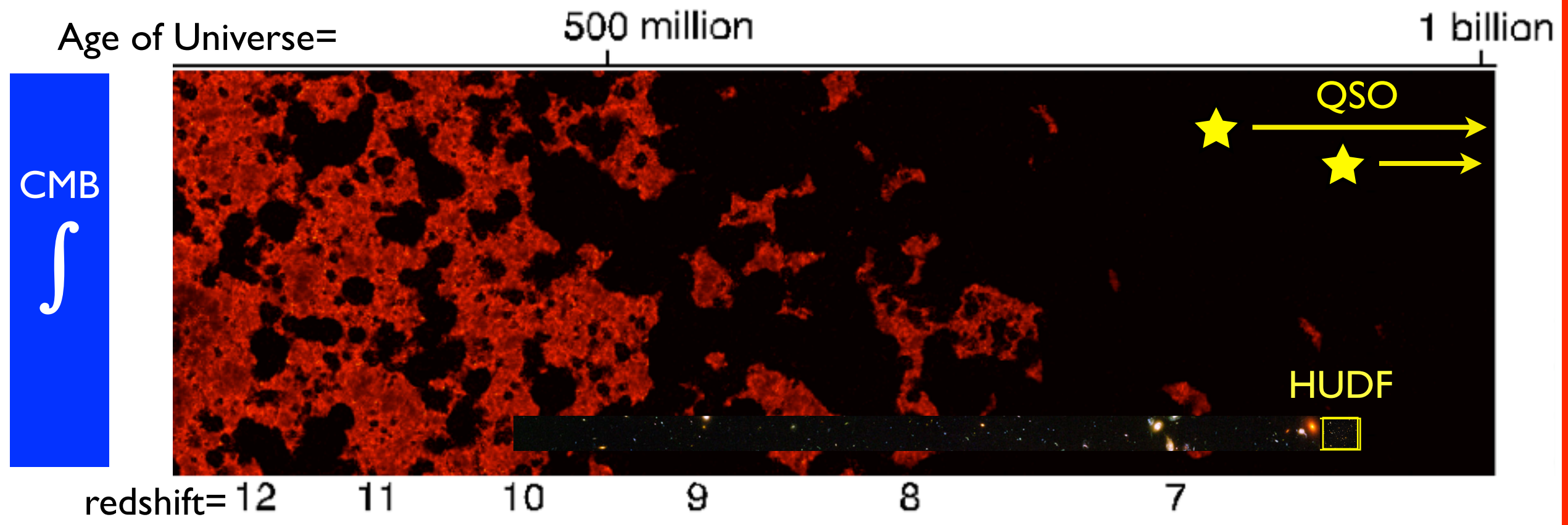
# DARE







# More needed...



Existing observations leaves much unanswered:

- 1) Lyman-alpha forest: end point  $z > 6.5$
- 2) CMB optical depth: mid point  $z \sim 11$
- 3) kSZ amplitude: duration  $z < 4.4$  ?

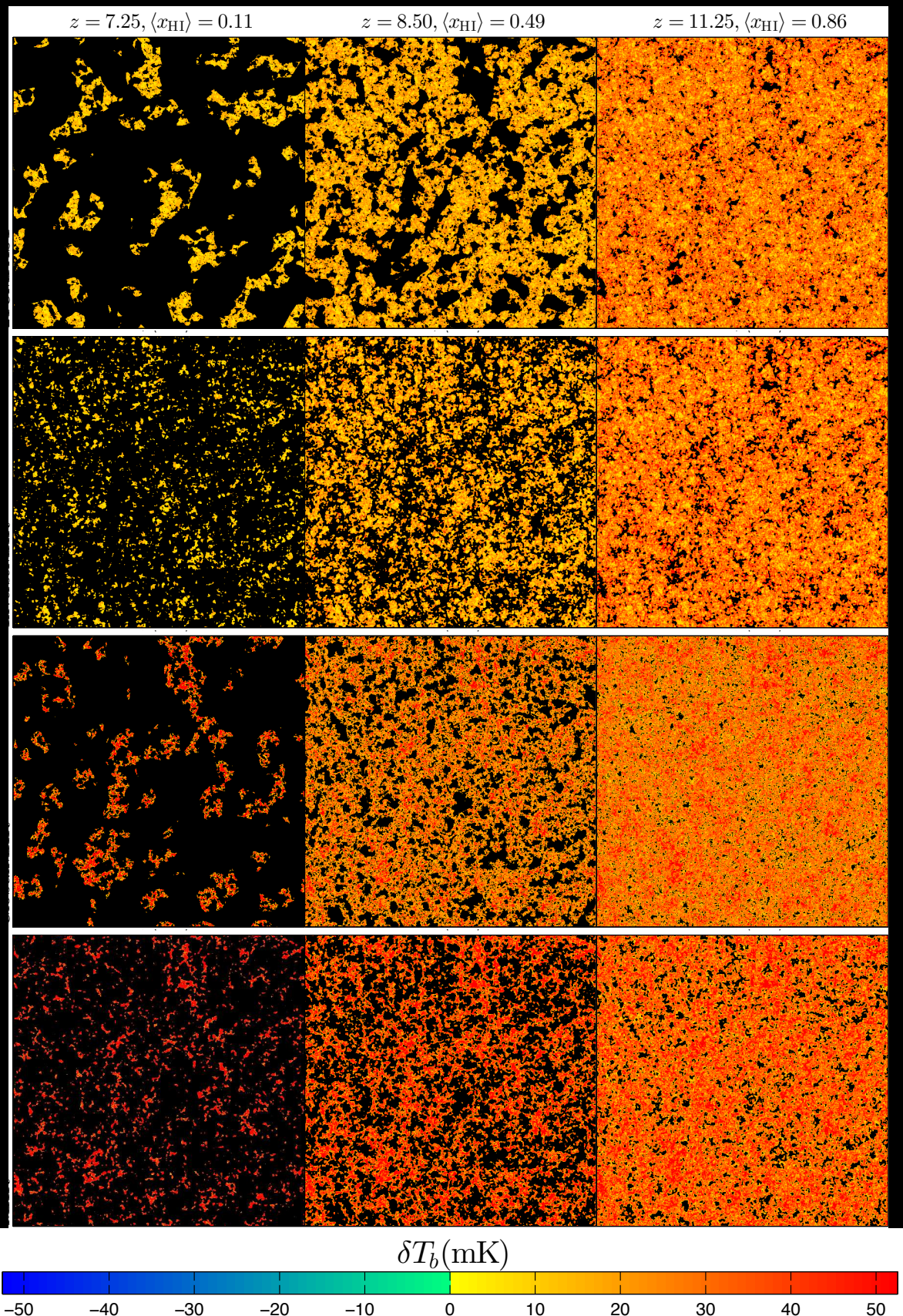
HST probes skewer much smaller than scale of ionized regions + only brightest sources

Large galaxy samples with LAE surveys or Euclid possible to  $z \sim 8$

Fundamental need for new types of observation to understand details of reionization

# Simulation of galaxy formation





We could imagine that light moved more slowly  
say 1 km/year

Then 18000 km to Wellington, NZ would take 18000 years  
we'd see back to the beginning of human history

and you'd have to wait ~18 hours after hitting  
a lightswitch to see the light



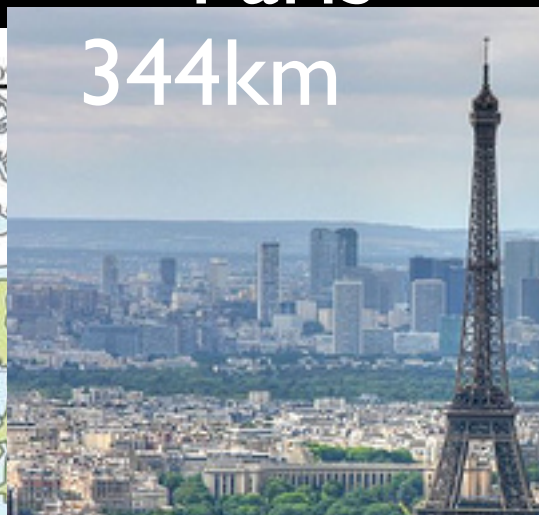
London

0km



Paris

344km



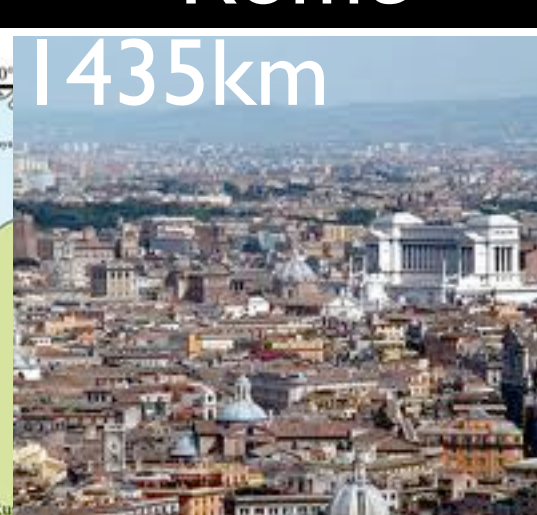
Berlin

932km



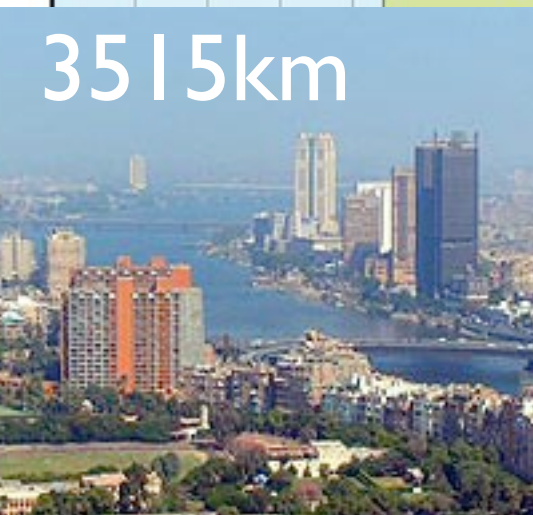
Rome

1435km



Cairo

3515km



Baghdad

4097km



Delhi

6717km



Sydney

17012km



Wellington

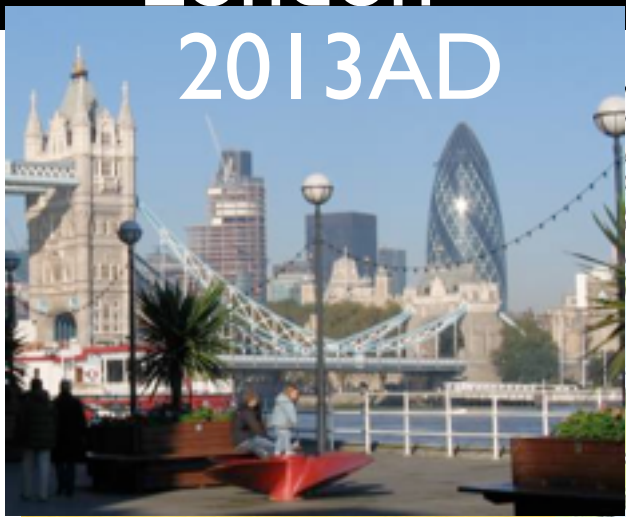
18834km





London

2013AD



Paris

1669AD



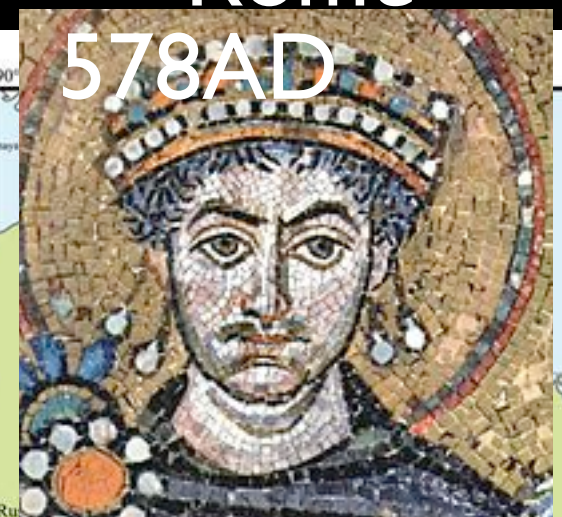
Berlin

1081AD



Rome

578AD



Cairo



Baghdad



2084BC

Delhi



4500BC

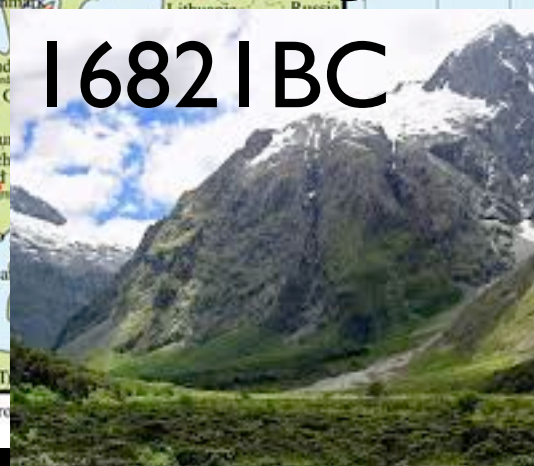
Sydney



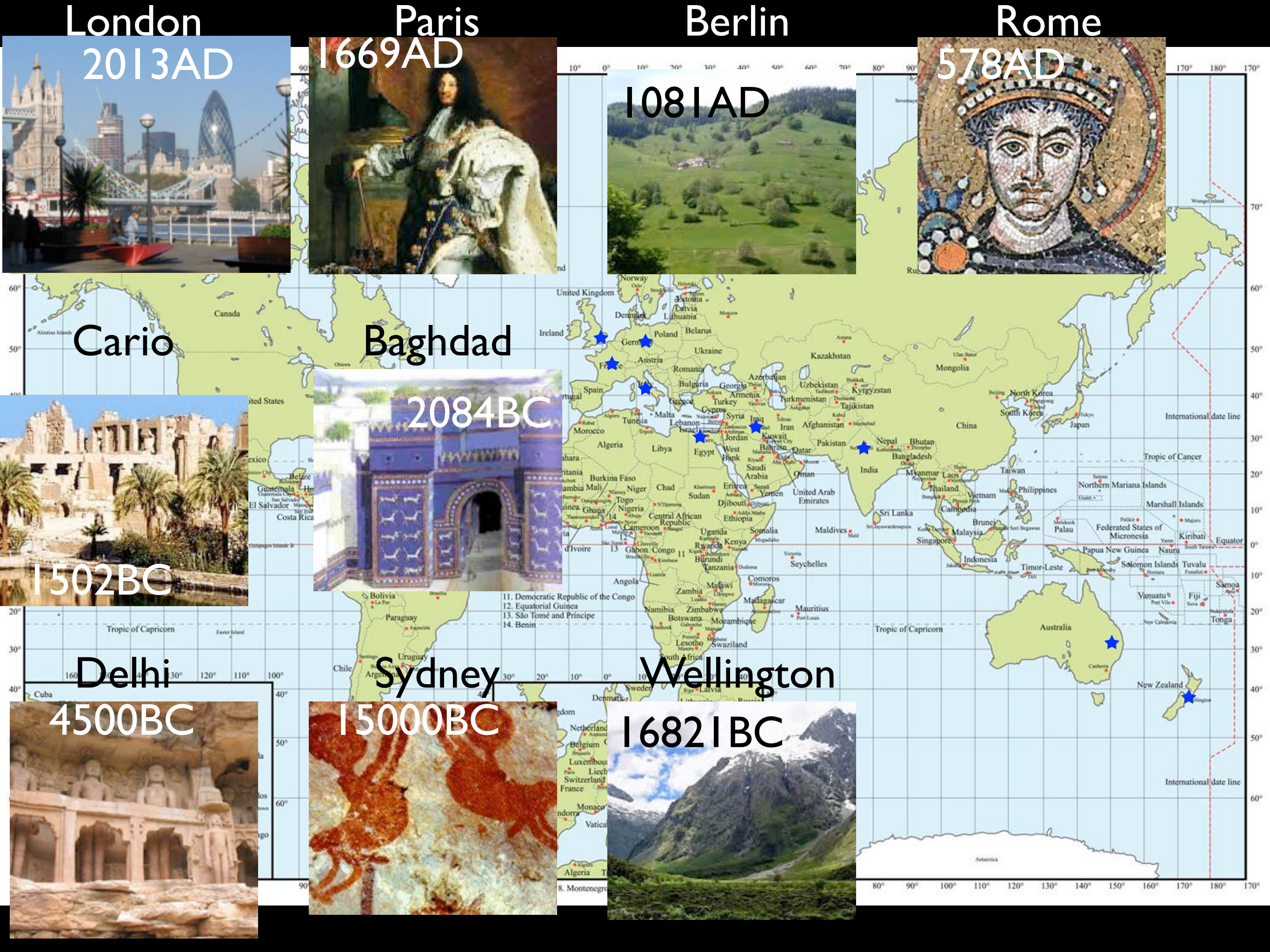
15000BC



Wellington



16821BC





# Stellar Archeology

- Our galaxy has formed via mergers - “hierarchical structure formation” - smaller units form first then merger to form larger ones
- Some components were galaxies that formed long ago
- Low mass stars have lifetime > age universe so may be still around today
- Search for metal-poor and metal-free stars in our Milky way

## LETTER

doi:10.1038/nature12990

**A single low-energy, iron-poor supernova  
as the source of metals in the star SMSS  
J031300.36–670839.3**

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